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# Nonlinear thin-walled beams with rectangular cross-section - Part II 

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#### Abstract

In this paper we report the second part of our results concerning the rigorous derivation of a hierarchy of one-dimensional models for thinwalled beams with rectangular cross-section. Denoting by $h$ and $\delta_{h} \ll h$ the length of the sides of the cross-section of the beam, we analyse the limit behaviour of a non-linear elastic energy which scales as $\varepsilon_{h}^{2}$ when $\varepsilon_{h} / \delta_{h} \rightarrow 0$.


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## 1 Introduction

The purpose of this paper is to continue the rigorous derivation started in [8] of a hierarchy of one-dimensional models for thin-walled beams. As explained in Part I, geometrically, a thin-walled beam is a slender structural element whose length is much larger than the diameter of the cross-section which, on its hand, is larger than the thickness of the thin wall. To model it, we consider a beam of length $\ell$ with a rectangular cross-section of sides $h$ and $\delta_{h}$ with

$$
h \rightarrow 0 \quad \text { and } \quad \frac{\delta_{h}}{h} \xrightarrow{h \rightarrow 0} 0 .
$$

After rescaling the domain the elastic energy rewrites as

$$
I^{h}(y)=\int_{\Omega} W\left(\nabla_{h} y(x)\right) d x, \quad \text { with } \quad \nabla_{h} y=\left(y, 1, \frac{y, 2}{h}, \frac{y, 3}{\delta_{h}}\right)
$$

[^0]where $W$ denotes the elastic energy density of the material, while $y$ and $\nabla_{h} y$ denote, respectively, the deformation and the rescaled deformation gradient.

We let $\left(y^{h}\right)$ be a sequence of deformations for which the energy scales as $\varepsilon_{h}^{2}$, where $\left(\varepsilon_{h}\right)$ is a sequence of positive numbers; more precisely, we assume that

$$
\begin{equation*}
I^{h}\left(y^{h}\right) \leq C \varepsilon_{h}^{2} \tag{1}
\end{equation*}
$$

and we study the $\Gamma$-limit of the sequence of functionals $I^{h} / \varepsilon_{h}^{2}$. The expression of the $\Gamma$-limit depends on the behaviour of $\varepsilon_{h}$ with respect to the intrinsic scale $\delta_{h}$. More precisely, we identify three main regimes:

- subcritical: $\frac{\delta_{h}}{\varepsilon_{h}} \xrightarrow{h \rightarrow 0} 0$;
- critical: $\frac{\delta_{h}}{\varepsilon_{h}} \xrightarrow{h \rightarrow 0} 1$;
- supercritical: $\frac{\delta_{h}}{\varepsilon_{h}} \xrightarrow{h \rightarrow 0}+\infty$.

The subcritical and the critical regimes have been studied in [8]. In this paper we focus on the supercritical case.

Assuming $\varepsilon_{h} / \delta_{h} \rightarrow 0$, we first show that, if a sequence of deformations ( $y^{h}$ ) satisfies (1), then the rescaled gradients $\nabla_{h} y^{h}$ must converge, as $h \rightarrow 0$, to a constant rotation (Lemma 3.1), which can be assumed to coincide with the identity, up to an orthonormal change of coordinates. Therefore, we expect to have linearization effects in the limiting energy. For this reason we introduce the sequence of displacements $\left(u^{h}\right)$ and of twist functions $\left(\vartheta^{h}\right)$ associated with $\left(y^{h}\right)$ and study their compactness properties (see Lemma 3.7). This part of the proof deeply relies on the rigidity estimate obtained by Friesecke, James, and Müller [9].

We then show that the $\Gamma$-limit of $I^{h} / \varepsilon_{h}^{2}$, as $h \rightarrow 0$, can be expressed in terms of the limit displacement $u$ and of the limit twist function $\vartheta$, and depends on the existence and on the value of the following limit:

$$
r:=\lim _{h \rightarrow 0} \frac{\varepsilon_{h}}{\delta_{h}^{2}}
$$

We distinguish the three regimes $r=0, r=+\infty$, and $r \sim 1$. By a rescaling of the cross-section, the last one can be reduced to the case $r=1$. If $r \in\{0,1\}$, we first prove that the limit displacement $u$ must belong to the set $\mathcal{A}^{B N}$ of Bernoulli-Navier displacements (see Definition 3.6). Moreover, in Theorems 3.11 and 3.14 we show that for these values of $r$ the $\Gamma$-limit of $I^{h} / \varepsilon_{h}^{2}$ is the functional $I^{r}: \mathcal{A}^{B N} \times W^{1,2}(0, \ell) \rightarrow[0,+\infty)$ defined by
$I^{r}(u, \vartheta):=\frac{1}{24} \int_{0}^{\ell} Q_{2}\left(\xi_{3}^{\prime \prime}, \vartheta^{\prime}\right) d x_{1}+\frac{1}{2} \int_{0}^{\ell} E\left(\xi_{1}^{\prime}+\frac{r}{2}\left(\xi_{3}^{\prime}\right)^{2}\right)^{2} d x_{1}+\frac{1}{24} \int_{0}^{\ell} E\left(\xi_{2}^{\prime \prime}\right)^{2} d x_{1}$
for every $(u, \vartheta) \in \mathcal{A}^{B N} \times W^{1,2}(0, \ell)$. Here the functions $\xi_{1} \in W^{1,2}(0, \ell), \xi_{2}, \xi_{3} \in$ $W^{2,2}(0, \ell)$ are such that

$$
u_{1}(x)=\xi_{1}\left(x_{1}\right)-x_{2} \xi_{2}^{\prime}\left(x_{1}\right)-x_{3} \xi_{3}^{\prime}\left(x_{1}\right), \quad u_{2}(x)=\xi_{2}\left(x_{1}\right), \quad u_{3}(x)=\xi_{3}\left(x_{1}\right)
$$

for a.e. $x \in \Omega$. The density function $Q_{2}$ is a positive definite quadratic form, while $E$ is a positive constant, and they both can be easily computed from the knowledge of $W$ (see (6) and (7)). If the beam is made of an isotropic material, the constant $E$ coincides with the Young modulus of the material.

If, instead, $r=+\infty$, we prove that the limit displacement $u$ must have the following structure: for a.e. $x \in \Omega$

$$
u_{1}(x)=\xi_{1}\left(x_{1}\right), \quad u_{2}(x)=0, \quad u_{3}(x)=\xi_{3}\left(x_{1}\right)
$$

with $\xi_{1}, \xi_{3} \in W^{2,2}(0, \ell)$ satisfying

$$
\xi_{1}^{\prime}=-\frac{1}{2}\left(\xi_{3}^{\prime}\right)^{2}
$$

We denote by $\mathcal{A}^{\infty}$ the set of all displacements in $W^{2,2}\left(\Omega ; \mathbb{R}^{3}\right)$ satisfying these conditions. Assuming in addition that

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{h^{2} \varepsilon_{h}}{\delta_{h}^{2}}=0 \tag{2}
\end{equation*}
$$

in Theorems 3.11 and 3.12 we show that, for $r=+\infty$, the $\Gamma$-limit of $I^{h} / \varepsilon_{h}^{2}$ is given by the functional

$$
I^{\infty}(u, \vartheta):=\frac{1}{24} \int_{0}^{\ell} Q_{2}\left(u_{3}^{\prime \prime}, \vartheta^{\prime}\right) d x_{1}
$$

for every $(u, \vartheta) \in \mathcal{A}^{\infty} \times W^{1,2}(0, \ell)$.
Assumption (2) is crucial in the construction of the recovery sequence. Heuristically, it allows us to stretch the mid-plane, i.e., the $x_{1} x_{2}$-plane, by deformations of order $\varepsilon_{h} /\left(\delta_{h} / h\right)^{2}$. When $\lim _{h \rightarrow 0} \varepsilon_{h} /\left(\delta_{h} / h\right)^{2} \neq 0$ the mid-plane must undergo a deformation which is very close to an infinitesimal isometry. For this reason we conjecture that, in this range, the $\Gamma$-limit should coincide with the $\Gamma$-limit of the geometrically linear Kirchhoff functional for a rectangular plate (see [10]), representing the mid-plane of the beam, when the length of one of the two sides approaches zero.
$\Gamma$-convergence results for thin-walled beams were obtained within the theory of linear elasticity in $[5,6,7]$, while $\Gamma$-convergence results for beams within the nonlinear framework were deduced in $[1,12,13,14,15]$.

The paper is organized as follows. In Section 2 we recall the setting of the problem and some preliminary results. Section 3 is devoted to the discussion of the supercritical case. Finally, in Section 4 we introduce applied loads and prove convergence of minimizers.

The notation is the same adopted in Part I of the present paper, to which we refer for details.

## 2 Setting of the problem and preliminaries

Let

$$
\Omega_{h}:=(0, \ell) \times \omega_{h} \subset \mathbb{R}^{3},
$$

where

$$
\omega_{h}:=\left\{\left(z_{2}, z_{3}\right):\left|z_{2}\right|<h / 2,\left|z_{3}\right|<\delta_{h} / 2\right\} \subset \mathbb{R}^{2}
$$

with $h>0, \delta_{1}:=1$ and

$$
\lim _{h \rightarrow 0} \delta_{h} / h=0 .
$$

Henceforth we shall refer to $\Omega_{h}$ as the reference configuration of the body and denote the elastic energy associated with a deformation $v: \Omega_{h} \rightarrow \mathbb{R}^{3}$ by

$$
E^{h}(v):=\int_{\Omega_{h}} W(\nabla v(z)) d z .
$$

We assume that the stored energy density $W: \mathbb{R}^{3 \times 3} \rightarrow[0,+\infty]$ satisfies the following assumptions:

1. $W \in C^{0}\left(\mathbb{R}^{3 \times 3}\right), W$ is of class $C^{2}$ in a neighborhood of $S O(3)$;
2. $W$ is frame indifferent, i.e., $W(F)=W(R F)$ for every $F \in \mathbb{R}^{3 \times 3}$ and $R \in S O(3)$;
3. $W(F) \geq C \operatorname{dist}^{2}(F, S O(3)), C>0 ; W(F)=0$ if $F \in S O(3)$.

A key role will be played by the following quadratic form:

$$
\begin{equation*}
Q_{3}(F):=\frac{\partial^{2} W}{\partial F^{2}}(I)(F, F)=\sum_{i, j, k, l=1}^{3} \frac{\partial^{2} W}{\partial F_{i j} \partial F_{k l}}(I) F_{i j} F_{k l}, \quad F \in \mathbb{R}^{3 \times 3} . \tag{3}
\end{equation*}
$$

In view of 3 this form is positive semi-definite and hence convex. Moreover, by 1 and 2 we have that (see, e.g., [11, Section 29])

$$
\begin{equation*}
Q_{3}(F)=Q_{3}\left(\frac{F+F^{T}}{2}\right) . \tag{4}
\end{equation*}
$$

In the special case when the energy density $W$ is isotropic, that is, $W(R F Q)=$ $W(F)$ for all $F \in \mathbb{R}^{3 \times 3}$ and $R, Q \in S O(3)$, then it turns out that

$$
\begin{equation*}
Q_{3}(F)=2 \mu|e|^{2}+\lambda(\operatorname{tr} e)^{2}, \quad e=\frac{F+F^{T}}{2} \tag{5}
\end{equation*}
$$

for some $\lambda, \mu \in \mathbb{R}$.
The limit problems will be stated in terms of the density function

$$
\begin{equation*}
Q_{2}(\alpha, \beta):=\min \left\{Q_{3}(A): A \in \mathbb{R}^{3 \times 3}, A^{T}=A, A_{11}=\alpha, A_{12}=\beta\right\}, \tag{6}
\end{equation*}
$$

and of the constant

$$
\begin{equation*}
E:=\min \left\{Q_{2}(1, \beta): \beta \in \mathbb{R}\right\}=\min \left\{Q_{3}(A): A \in \mathbb{R}^{3 \times 3}, A^{T}=A, A_{11}=1\right\} \tag{7}
\end{equation*}
$$

Let us remark that $Q_{2}$ is a positive definite quadratic form and $E>0$. Moreover, in the isotropic case where $Q_{3}$ takes the form (5), a simple computation shows that

$$
Q_{2}(\alpha, \beta)=4 \mu \beta^{2}+E \alpha^{2}
$$

and $E=\mu \frac{2 \mu+3 \lambda}{\mu+\lambda}$ is the Young modulus of the material.
To state our results it is convenient to stretch the domain $\Omega_{h}$ along the transverse directions $z_{2}$ and $z_{3}$ in a way that the transformed domain does not depend on $h$. Let us therefore set $\omega:=\omega_{1}, \Omega:=\Omega_{1}$, and let

$$
p_{h}: \Omega \rightarrow \Omega_{h}
$$

be defined by

$$
\begin{equation*}
p_{h}(x)=p_{h}\left(x_{1}, x_{2}, x_{3}\right)=\left(x_{1}, h x_{2}, \delta_{h} x_{3}\right) \tag{8}
\end{equation*}
$$

Let us consider the following $3 \times 3$ matrix

$$
\begin{equation*}
\nabla_{h} y:=\left(y,_{1}, \frac{y, 2}{h}, \frac{y, 3}{\delta_{h}}\right) \tag{9}
\end{equation*}
$$

where $y,{ }_{i}$ denotes the column vector of the partial derivatives of $y$ with respect to $x_{i}, i=1,2,3$. Then we can consider the rescaled energy $I^{h}: W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right) \rightarrow$ $[0,+\infty]$ defined by $I^{h}(y):=\frac{1}{h \delta_{h}} E^{h}\left(y \circ p_{h}^{-1}\right)$, i.e.,

$$
I^{h}(y)=\int_{\Omega} W\left(\nabla_{h} y(x)\right) d x
$$

for every $y \in W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$.
Throughout the rest of the paper $\left(\varepsilon_{h}\right)$ will denote a sequence of strictly positive real numbers. We conclude the section by recalling a result proven in [8, Theorem 3.2 and Lemma 3.3] and concerning some general compactness properties for sequences of deformations with equibounded energy.
Lemma 2.1 Let $\left(y^{h}\right)$ be a sequence in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ such that

$$
\begin{equation*}
\left(\int_{\Omega} \operatorname{dist}^{2}\left(\nabla_{h} y^{h}, S O(3)\right) d x\right)^{\frac{1}{2}} \leq C \varepsilon_{h} \tag{10}
\end{equation*}
$$

for every $h>0$. Then, there exist two sequences $R^{h}:(0, \ell) \times(-1 / 2,1 / 2) \rightarrow$ $S O(3)$ and

$$
\tilde{R}^{h} \in C^{\infty}\left((0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right) ; \mathbb{R}^{3 \times 3}\right)
$$

such that

1. $\left\|\tilde{R}^{h}-R^{h}\right\|_{L^{2}} \leq C \varepsilon_{h}, \quad\left\|\tilde{R}^{h}-R^{h}\right\|_{L^{\infty}} \leq C h^{1 / 2} \varepsilon_{h} / \delta_{h}$,
2. $\left\|\nabla_{h} y^{h}-\tilde{R}^{h}\right\|_{L^{2}} \leq C \varepsilon_{h}$,
3. $\left\|\tilde{R}_{, 1}^{h}\right\|_{L^{2}} \leq C \varepsilon_{h} / \delta_{h}, \quad\left\|\tilde{R}_{, 2}^{h}\right\|_{L^{2}} \leq C h \varepsilon_{h} / \delta_{h}$,
where the constant $C$ may change from line to line. Moreover, if in addition $h^{1 / 2} \varepsilon_{h} / \delta_{h} \rightarrow 0$, then we can take
4. $\tilde{R}^{h}\left(x_{1}, x_{2}\right) \in S O(3)$ for every $\left(x_{1}, x_{2}\right) \in(0, \ell) \times(-1 / 2,1 / 2)$ and every $h>0$.

## 3 The supercritical case

This section is devoted to the study of the asymptotic behaviour of a sequence of deformations $\left(y^{h}\right) \subset W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ satisfying

$$
\begin{equation*}
I^{h}\left(y^{h}\right)=\int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x \leq C \varepsilon_{h}^{2} \tag{11}
\end{equation*}
$$

for every $h>0$, where

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{\varepsilon_{h}}{\delta_{h}}=0 \tag{12}
\end{equation*}
$$

Under these assumptions, properties 3 and 4 of Lemma 2.1 imply that the sequence $\tilde{R}^{h}$ converges weakly in $W^{1,2}$ to a constant rotation $R$. In the next lemma we introduce suitable rotations and translations of the coordinate system in such a way to deal with a limit rotation equal to the identity.

Lemma 3.1 Let $\left(y^{h}\right)$ be a sequence in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ satisfying (10) and let ( $\left.\tilde{R}^{h}\right)$ be the sequence constructed in Lemma 2.1. Then, there exists a sequence of constant rotations $Q^{h} \in S O(3)$ such that, setting $\bar{R}^{h}=Q^{h^{T}} \tilde{R}^{h}$ and $\bar{y}^{h}=$ $Q^{h^{T}} y^{h}-c^{h}$, where $c^{h}$ is any constant, we have

1. $\left\|\nabla_{h} \bar{y}^{h}-\bar{R}^{h}\right\|_{L^{2}} \leq C \varepsilon_{h}$,
2. $\left\|\bar{R}^{h}{ }_{, 1}\right\|_{L^{2}} \leq C \frac{\varepsilon_{h}}{\delta_{h}}, \quad\left\|\bar{R}^{h}{ }_{, 2}\right\|_{L^{2}} \leq C h \frac{\varepsilon_{h}}{\delta_{h}}$,
3. $\left\|\bar{R}^{h}-I\right\|_{L^{2}} \leq C \frac{\varepsilon_{h}}{\delta_{h}}$,
where the constant $C$ may change from line to line. Moreover, if in addition (12) holds, then for every $h$ small enough we can take
4. $\bar{R}^{h}\left(x_{1}, x_{2}\right) \in S O(3)$ for every $\left(x_{1}, x_{2}\right) \in(0, \ell) \times(-1 / 2,1 / 2)$
and

$$
\begin{equation*}
\int_{\Omega}\left(\nabla_{h} \bar{y}^{h}-\nabla_{h} \bar{y}^{h^{T}}\right) d x=0 \tag{13}
\end{equation*}
$$

Proof. By Sobolev-Poincaré inequality there exist some constant matrices $\tilde{Q}^{h} \in \mathbb{R}^{3 \times 3}$ such that

$$
\begin{equation*}
\left\|\tilde{R}^{h}-\tilde{Q}^{h}\right\|_{L^{2}} \leq C\left\|\nabla \tilde{R}^{h}\right\|_{L^{2}} \leq C \frac{\varepsilon_{h}}{\delta_{h}} \tag{14}
\end{equation*}
$$

where the last inequality follows from property 3 of Lemma 2.1. Let $R^{h}$ be the sequence of approximating rotations constructed in Lemma 2.1. The first inequality in 1 of that lemma and (14) yield

$$
\left\|R^{h}-\tilde{Q}^{h}\right\|_{L^{2}} \leq C \frac{\varepsilon_{h}}{\delta_{h}}
$$

and since $R^{h} \in S O(3)$, this implies that

$$
\operatorname{dist}\left(\tilde{Q}^{h}, S O(3)\right) \leq C \frac{\varepsilon_{h}}{\delta_{h}}
$$

Thus, there exists $\hat{Q}^{h} \in S O(3)$ such that

$$
\left|\hat{Q}^{h}-\tilde{Q}^{h}\right| \leq C \frac{\varepsilon_{h}}{\delta_{h}}
$$

Setting $\hat{R}^{h}:=\hat{Q}^{h^{T}} \tilde{R}^{h}$ and using (14), we obtain

$$
\left\|\hat{R}^{h}-I\right\|_{L^{2}}=\left\|\tilde{R}^{h}-\hat{Q}^{h}\right\|_{L^{2}} \leq\left\|\tilde{R}^{h}-\tilde{Q}^{h}\right\|_{L^{2}}+\left\|\tilde{Q}^{h}-\hat{Q}^{h}\right\|_{L^{2}} \leq C \frac{\varepsilon_{h}}{\delta_{h}}
$$

that is, property 3 of the statement for the sequence $\hat{R}^{h}$. Moreover, setting $\hat{y}^{h}:=\hat{Q}^{h^{T}} y^{h}-\hat{c}^{h}$, where $\hat{c}^{h}$ is any constant, we deduce properties 1 and 2 of the statement for $\left(\hat{y}^{h}\right)$ and $\left(\hat{R}^{h}\right)$ from properties 2 and 3 of Lemma 2.1.

Assume now (12). Property 4 of the statement for $\left(\hat{R}^{h}\right)$ follows immediately from 4 of Lemma 2.1. In order to satisfy also (13) we need to modify the constructed sequences. Let

$$
F^{h}:=f_{\Omega} \nabla_{h} \hat{y}^{h} d x
$$

Then, from properties 1 and 3 for the sequences with an over-hat, we have

$$
\begin{align*}
\left|F^{h}-I\right| & \leq f_{\Omega}\left|\nabla_{h} \hat{y}^{h}-I\right| d x \leq C\left\|\nabla_{h} \hat{y}^{h}-I\right\|_{L^{2}} \\
& \leq C\left(\left\|\nabla_{h} \hat{y}^{h}-\hat{R}^{h}\right\|_{L^{2}}+\left\|\hat{R}^{h}-I\right\|_{L^{2}}\right) \leq C \frac{\varepsilon_{h}}{\delta_{h}} \tag{15}
\end{align*}
$$

By (12) this implies, in particular, that det $F^{h}>0$ for $h$ sufficiently small. Thus, by the polar decomposition theorem, there exist $P^{h} \in S O(3)$ and a positive symmetric matrix $U^{h}$ such that $F^{h}=P^{h} U^{h}$. Since $\left|U^{h}-I\right|=\operatorname{dist}\left(F^{h}, S O(3)\right) \leq$ $\left|F^{h}-I\right|$, we have

$$
\begin{equation*}
\left|P^{h}-I\right| \leq\left|P^{h}-F^{h}\right|+\left|F^{h}-I\right|=\left|U^{h}-I\right|+\left|F^{h}-I\right| \leq C \frac{\varepsilon_{h}}{\delta_{h}} \tag{16}
\end{equation*}
$$

where we have used (15). We claim that

$$
Q^{h}:=\hat{Q}^{h} P^{h}, \quad \bar{R}^{h}:=P^{h^{T}} \hat{R}^{h}, \quad \bar{y}^{h}:=P^{h^{T}} \hat{y}^{h}=Q^{h^{T}} y^{h}-c^{h}
$$

satisfy properties $1-4$ and (13) of the lemma. Indeed, conditions 1, 2, and 4 are immediate, while 3 follows from (16), since we have

$$
\left\|\bar{R}^{h}-I\right\|_{L^{2}} \leq\left\|\bar{R}^{h}-\hat{R}^{h}\right\|_{L^{2}}+\left\|\hat{R}^{h}-I\right\|_{L^{2}}=\left\|P^{h}-I\right\|_{L^{2}}+\left\|\hat{R}^{h}-I\right\|_{L^{2}}
$$

Finally, since $\nabla_{h} \bar{y}^{h}=P^{h^{T}} \nabla_{h} \hat{y}^{h}$, we have

$$
\begin{aligned}
\int_{\Omega}\left(\nabla_{h} \bar{y}^{h}-\nabla_{h} \bar{y}^{h^{T}}\right) d x & =|\Omega|\left(P^{h^{T}} F^{h}-\left(P^{h^{T}} F^{h}\right)^{T}\right) \\
& =|\Omega|\left(U^{h}-U^{h^{T}}\right)=0
\end{aligned}
$$

hence also (13) is satisfied.

Remark 3.2 Since the energy density $W$ is frame indifferent, the energy $I^{h}$ on a deformation $y$ does not change if a rigid motion is superimposed to $y$; therefore, a sequence of deformations $\left(y^{h}\right)$ satisfying (11) is not, in general, bounded in any reasonable space. In Lemma 3.1 to obtain bounds we have superimposed an appropriate rigid motion $r^{h}(x):=Q^{h^{T}} x-c^{h}$ to each deformation $y^{h}$. The motion $r^{h}$ is not uniquely determined; indeed, if we replace $Q^{h}$ by $Q^{h} \exp \left(\frac{\varepsilon_{h}}{\delta_{h}} K\right)$, then properties 1, 2, 3 are still satisfied and condition 4 may be obtained arguing as in the proof of the lemma. However, one can easily show that if $\left(Q^{h}\right)$ and $\left(\tilde{Q}^{h}\right)$ are two sequences of constant rotations for which the lemma is true, then $\left|Q^{h}-\tilde{Q}^{h}\right| \leq C \frac{\varepsilon_{h}}{\delta_{h}}$.

In the next lemma we study the implications of the bounds obtained in Lemma 3.1.

Lemma 3.3 Assume (12). Let $\left(y^{h}\right)$ be a sequence in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ satisfying (10) and let $\left(\bar{R}^{h}\right)$ be the sequence constructed in Lemma 3.1. Then there exist three tensor fields $A \in W^{1,2}\left((0, \ell) ; \mathbb{R}^{3 \times 3}\right), B \in L^{2}\left((0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right) ; \mathbb{R}^{3 \times 3}\right)$, and $G \in L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$, with $A$ and $B$ skew-symmetric, such that, up to subsequences,

1. $A^{h}:=\frac{\bar{R}^{h}-I}{\varepsilon_{h} / \delta_{h}} \rightharpoonup A$ in $W^{1,2}\left((0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right) ; \mathbb{R}^{3 \times 3}\right)$,
2. $\operatorname{sym} \frac{\bar{R}^{h}-I}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}} \rightarrow \frac{A^{2}}{2}$ in $L^{2}\left((0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right) ; \mathbb{R}^{3 \times 3}\right)$,
3. $B^{h}:=\frac{\bar{R}^{h}, 2}{h \varepsilon_{h} / \delta_{h}} \rightharpoonup B$ in $L^{2}\left((0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right) ; \mathbb{R}^{3 \times 3}\right)$,
4. $G^{h}:=\frac{\bar{R}^{h^{T}} \nabla_{h} \bar{y}^{h}-I}{\varepsilon_{h}} \rightharpoonup G$ in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$.

Moreover, we have
5. $A_{, 1} e_{2}=B e_{1}$, hence $A_{12,1}=B_{12}=0$ and $A_{23,1}=B_{13}$ a.e. in $(0, \ell) \times$ $\left(-\frac{1}{2}, \frac{1}{2}\right)$,
6. $G(x) e_{1}=x_{3} A_{1}\left(x_{1}\right) e_{3}+\tilde{G}\left(x_{1}, x_{2}\right) e_{1}$ for a.e. $x \in \Omega$,
7. $G(x) e_{2}=x_{3} B\left(x_{1}, x_{2}\right) e_{3}+\tilde{G}\left(x_{1}, x_{2}\right) e_{2}$ for a.e. $x \in \Omega$,
for a suitable $\tilde{G} \in L^{2}\left((0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right) ; \mathbb{R}^{3 \times 3}\right)$.
Proof. By 3 of Lemma 3.1 the sequence $A^{h}$ is bounded in $L^{2}$, hence it admits a subsequence which converges weakly in $L^{2}$. Let $A$ denote this weak limit. By 2 of Lemma 3.1 we have that $A_{, 2}^{h} \rightarrow 0$ in $L^{2}$, while the derivative with respect to $x_{1}$ is bounded in $L^{2}$. This implies that, up to subsequences, $A^{h} \rightharpoonup A$ weakly in $W^{1,2}$ and that the limit $A$ is independent of $x_{2}$. Since $\bar{R}^{h} \in S O(3)$, we have

$$
\begin{equation*}
A^{h}+A^{h^{T}}=-\frac{\varepsilon_{h}}{\delta_{h}} A^{h^{T}} A^{h}, \tag{17}
\end{equation*}
$$

and passing to the limit as $h \rightarrow 0$, we obtain that $A+A^{T}=0$.
We now prove 2. By (17) we have that

$$
\operatorname{sym} \frac{\bar{R}^{h}-I}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}}=\operatorname{sym} \frac{A^{h}}{\varepsilon_{h} / \delta_{h}}=-\frac{A^{h^{T}} A^{h}}{2} .
$$

The claim now follows from 1 , the compact embedding theorem, and the fact that $A$ is skew-symmetric.

Let us prove 3. The weak convergence of a subsequence of $B^{h}$ follows from the second estimate in 2 of Lemma 3.1. Let us call $B$ its weak limit. Since

$$
0=\left(\frac{\bar{R}^{h^{T}} \bar{R}^{h}}{h \varepsilon_{h} / \delta_{h}}\right)_{, 2}=\frac{\bar{R}_{, 2}^{h^{T}}}{h \varepsilon_{h} / \delta_{h}} \bar{R}^{h}+\bar{R}^{h^{T}} \frac{\bar{R}_{, 2}^{h}}{h \varepsilon_{h} / \delta_{h}},
$$

we deduce that $\operatorname{sym} B=0$ by passing to the limit and using 3 of Lemma 3.1.
Convergence 4 is an immediate consequence of 1 of Lemma 3.1.
Property 5 follows from the equality

$$
A_{, 1}^{h} e_{2}=\left(\frac{\bar{R}^{h} e_{2}-\bar{y}_{, 2}^{h} / h}{\varepsilon_{h} / \delta_{h}}\right)_{, 1}+\left(\frac{\bar{y}_{, 1}^{h}-\bar{R}^{h} e_{1}}{h \varepsilon_{h} / \delta_{h}}\right)_{, 2}+\frac{\bar{R}^{h}{ }_{, 2} e_{1}}{h \varepsilon_{h} / \delta_{h}} \quad \text { in } H^{-1}\left(\Omega ; \mathbb{R}^{3 \times 3}\right) .
$$

Indeed, by using 1 of Lemma 3.1 and 3 to pass to the limit we obtain $A_{, 1} e_{2}=$ $B e_{1}$. This rewrites as $A_{i 2,1}=B_{i 1}$ for $i=1,2,3$, from which the remaining relations in 5 follow by using the fact that $A$ and $B$ are skew-symmetric.

To prove 6 we note that in $H^{-1}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$ there holds

$$
\left(\bar{R}^{h} G^{h} e_{1}\right)_{, 3}=\left(\frac{\bar{y}_{3}^{h} / \delta_{h}-\bar{R}^{h} e_{3}}{\varepsilon_{h} / \delta_{h}}\right)_{, 1}+\left(\frac{\bar{R}^{h} e_{3}-e_{3}}{\varepsilon_{h} / \delta_{h}}\right)_{, 1}
$$

and, using 1 and 3 of Lemma 3.1, and 1 and 4 already proven, we find

$$
G_{, 3} e_{1}=A_{, 1} e_{3}
$$

from which we obtain 6. Equation 7 can be proven similarly.
Hereafter we assume that the following limit exists:

$$
\begin{equation*}
r:=\lim _{h \rightarrow 0} \frac{\varepsilon_{h}}{\delta_{h}^{2}} \tag{18}
\end{equation*}
$$

Without loss of generality, we may assume that $r \in\{0,1,+\infty\}$, by possibly changing the value of the constant $C$, appearing in (11), and by taking $\omega_{h}=$ $(-a h / 2, a h / 2) \times\left(-b \delta_{h} / 2, b \delta_{h} / 2\right)$ for appropriate constants $a$ and $b$.

In the next lemma we take a closer look at the rescaled displacement gradient.
Lemma 3.4 Under the same assumptions of Lemma 3.3, we have
i. the sequence $\left(\frac{\nabla_{h} \bar{y}^{h}-I}{\varepsilon_{h} / \delta_{h}}\right)$ admits a subsequence which converges to $A$ in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$,
where $A$ is the field introduced in Lemma 3.3. With $r$ defined as in (18), the following statements hold:
ii. if $r=+\infty$, then, up to extracting a subsequence, $\operatorname{sym}\left(\frac{\nabla_{h} \bar{y}^{h}-I}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}}\right) \rightarrow \frac{A^{2}}{2}$ in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$,
iii. if $r \in\{0,1\}$, then $\operatorname{sym}\left(\frac{\nabla_{h} \bar{y}^{h}-I}{\varepsilon_{h}}\right)$ is bounded in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$.

Proof. Statement $i$ follows by observing that

$$
\frac{\nabla_{h} \bar{y}^{h}-I}{\varepsilon_{h} / \delta_{h}}=\frac{\nabla_{h} \bar{y}^{h}-\bar{R}^{h}}{\varepsilon_{h}} \delta_{h}+\frac{\bar{R}^{h}-I}{\varepsilon_{h} / \delta_{h}}
$$

and using 1 of Lemma 3.1 and 1 of Lemma 3.3.
Assume now (18). Statement $i i$ follows from

$$
\begin{equation*}
\operatorname{sym}\left(\frac{\nabla_{h} \bar{y}^{h}-I}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}}\right)=\operatorname{sym}\left(\frac{\nabla_{h} \bar{y}^{h}-\bar{R}^{h}}{\varepsilon_{h}}\right) \frac{\delta_{h}^{2}}{\varepsilon_{h}}+\operatorname{sym}\left(\frac{\bar{R}^{h}-I}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}}\right) \tag{19}
\end{equation*}
$$

and by 1 of Lemma 3.1 and 2 of Lemma 3.3.
Similarly, for $i i i$ we have

$$
\begin{equation*}
\operatorname{sym}\left(\frac{\nabla_{h} \bar{y}^{h}-I}{\varepsilon_{h}}\right)=\operatorname{sym}\left(\frac{\nabla_{h} \bar{y}^{h}-\bar{R}^{h}}{\varepsilon_{h}}\right)+\operatorname{sym}\left(\frac{\bar{R}^{h}-I}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}}\right) \frac{\varepsilon_{h}}{\delta_{h}^{2}}, \tag{20}
\end{equation*}
$$

and again the claim follows from 1 of Lemma 3.1 and 2 of Lemma 3.3.
We now define two sets of displacements which will play a crucial role in what follows.

Definition 3.5 Let $\mathcal{A}^{\infty}$ be the class of all displacements $u \in W^{2,2}\left((0, \ell) ; \mathbb{R}^{3}\right)$ satisfying the following property: there exist $\xi_{1}, \xi_{3} \in W^{2,2}(0, \ell)$ such that

$$
u_{1}=\xi_{1}, \quad u_{2}=0, \quad u_{3}=\xi_{3}
$$

with

$$
\begin{equation*}
\xi_{1}^{\prime}=-\frac{1}{2}\left(\xi_{3}^{\prime}\right)^{2} \tag{21}
\end{equation*}
$$

Definition 3.6 Let $\mathcal{A}^{B N}$ be the set of (Bernoulli-Navier) displacements $u \in$ $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ satisfying the following condition: there exist $\xi_{1} \in W^{1,2}(0, \ell)$, $\xi_{2}, \xi_{3} \in W^{2,2}(0, \ell)$ such that

$$
u_{1}=\xi_{1}-x_{2} \xi_{2}^{\prime}-x_{3} \xi_{3}^{\prime}, \quad u_{2}=\xi_{2}, \quad u_{3}=\xi_{3}
$$

Let

$$
\mathcal{A}^{r}:= \begin{cases}\mathcal{A}^{\infty} & \text { if } r=+\infty \\ \mathcal{A}^{B N} & \text { if } r \in\{0,1\}\end{cases}
$$

In the next lemma we introduce the twist of the cross-section and we study its convergence together with the convergence of the displacements.

Lemma 3.7 Under the same assumptions of Lemma 3.3, let $\vartheta^{h}:(0, \ell) \rightarrow \mathbb{R}$ be defined by

$$
\vartheta^{h}:=\frac{1}{I_{0}} \frac{1}{\varepsilon_{h}} \int_{\omega}\left(\frac{\delta_{h}}{h} x_{2} \bar{y}_{3}^{h}-x_{3} \bar{y}_{2}^{h}\right) d x_{2} d x_{3}
$$

where

$$
I_{0}:=\int_{\omega}\left(x_{2}^{2}+x_{3}^{2}\right) d x_{2} d x_{3}=\frac{1}{6}
$$

Let $A, G$, and $\tilde{G}$ be the fields introduced in Lemma 3.3. Then

$$
\vartheta^{h} \rightharpoonup \vartheta:=A_{32} \text { in } W^{1,2}(0, \ell),
$$

and for a.e. $x \in \Omega$

$$
\begin{align*}
& G_{12}(x)=-x_{3} \vartheta^{\prime}\left(x_{1}\right)+\tilde{G}_{12}\left(x_{1}, x_{2}\right) \\
& G_{21}(x)=-x_{3} \vartheta^{\prime}\left(x_{1}\right)+\tilde{G}_{21}\left(x_{1}, x_{2}\right) \tag{22}
\end{align*}
$$

Let $r$ be as in (18). Then the following statements hold:
i. if $r=+\infty$, then, for a suitable choice of the constants $c^{h}$ in Lemma 3.1, the sequence of displacements $u^{h}: \Omega \rightarrow \mathbb{R}^{3}$ defined by

$$
\begin{align*}
u_{1}^{h} & :=\frac{\bar{y}_{1}^{h}-x_{1}}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}} \\
u_{2}^{h} & :=\frac{\bar{y}_{2}^{h}-h x_{2}}{\varepsilon_{h} / \delta_{h}}  \tag{23}\\
u_{3}^{h} & :=\frac{\bar{y}_{3}^{h}-\delta_{h} x_{3}}{\varepsilon_{h} / \delta_{h}}
\end{align*}
$$

admits a subsequence which converges in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ to a function $u \in$ $\mathcal{A}^{\infty}$. Moreover,

$$
\begin{equation*}
G_{11}(x)=-x_{3} \xi_{3}^{\prime \prime}\left(x_{1}\right)+\tilde{G}_{11}\left(x_{1}, x_{2}\right) \tag{24}
\end{equation*}
$$

for a.e. $x \in \Omega$.
ii. If $r \in\{0,1\}$, then, for a suitable choice of the constants $c^{h}$ in Lemma 3.1, the sequence of displacements $u^{h}: \Omega \rightarrow \mathbb{R}^{3}$ defined by

$$
\begin{align*}
u_{1}^{h} & :=\frac{\bar{y}_{1}^{h}-x_{1}}{\varepsilon_{h}}, \\
u_{2}^{h} & :=\frac{\bar{y}_{2}^{h}-h x_{2}}{\varepsilon_{h} / h}  \tag{25}\\
u_{3}^{h} & :=\frac{\bar{y}_{3}^{h}-\delta_{h} x_{3}}{\varepsilon_{h} / \delta_{h}},
\end{align*}
$$

admits a subsequence which converges weakly in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ to a function $u \in \mathcal{A}^{B N}$. Moreover,

$$
\begin{equation*}
G_{11}(x)=\xi_{1}^{\prime}\left(x_{1}\right)-x_{2} \xi_{2}^{\prime \prime}\left(x_{1}\right)-x_{3} \xi_{3}^{\prime \prime}\left(x_{1}\right)+\frac{r}{2}\left(\xi_{3}^{\prime}\left(x_{1}\right)\right)^{2} \tag{26}
\end{equation*}
$$

for a.e. $x \in \Omega$.
Proof. Since $A=A\left(x_{1}\right)$, the convergence in $i$ of Lemma 3.4 implies that

$$
\begin{aligned}
\frac{1}{h \varepsilon_{h} / \delta_{h}}\left(\bar{y}_{3}^{h}-\int_{-\frac{1}{2}}^{\frac{1}{2}} \bar{y}_{3}^{h} d x_{2}\right) & \rightarrow A_{32} x_{2} \quad \text { in } L^{2}(\Omega) \\
\frac{1}{\varepsilon_{h}}\left(\bar{y}_{2}^{h}-\int_{-\frac{1}{2}}^{\frac{1}{2}} \bar{y}_{2}^{h} d x_{3}\right) & \rightarrow-A_{32} x_{3} \quad \text { in } L^{2}(\Omega)
\end{aligned}
$$

Since $\vartheta^{h}$ can be written as

$$
\begin{aligned}
\vartheta^{h}= & \frac{1}{I_{0}} \frac{1}{h \varepsilon_{h} / \delta_{h}} \int_{\omega} x_{2}\left(\bar{y}_{3}^{h}-\int_{-\frac{1}{2}}^{\frac{1}{2}} \bar{y}_{3}^{h} d x_{2}\right) d x_{2} d x_{3} \\
& -\frac{1}{I_{0}} \frac{1}{\varepsilon_{h}} \int_{\omega} x_{3}\left(\bar{y}_{2}^{h}-\int_{-\frac{1}{2}}^{\frac{1}{2}} \bar{y}_{2}^{h} d x_{3}\right) d x_{2} d x_{3}
\end{aligned}
$$

it is clear that $\vartheta^{h}$ converges to $\vartheta:=A_{32}$ strongly in $L^{2}$. The convergence is actually weak in $W^{1,2}$, as $\left(\vartheta^{h}\right)^{\prime}$ is bounded in $L^{2}$. Indeed, using the fact that $\bar{R}^{h}$ is independent of $x_{3}$, we obtain

$$
\begin{aligned}
\left(\vartheta^{h}\right)^{\prime}= & \frac{1}{I_{0}} \frac{1}{h \varepsilon_{h} / \delta_{h}} \int_{\omega} x_{2}\left(\bar{y}_{3,1}^{h}-\bar{R}_{31}^{h}\right) d x_{2} d x_{3} \\
& +\frac{1}{I_{0}} \frac{1}{h \varepsilon_{h} / \delta_{h}} \int_{\omega} x_{2}\left(\bar{R}_{31}^{h}-\int_{-\frac{1}{2}}^{\frac{1}{2}} \bar{R}_{31}^{h} d x_{2}\right) d x_{2} d x_{3} \\
& -\frac{1}{I_{0}} \frac{1}{\varepsilon_{h}} \int_{\omega} x_{3}\left(\bar{y}_{2,1}^{h}-\bar{R}_{21}^{h}\right) d x_{2} d x_{3},
\end{aligned}
$$

where the first and the last term on the right-hand side are bounded in $L^{2}$ by 1 of Lemma 3.1, while the second term is bounded in $L^{2}$ by Poincaré-Wirtinger inequality and the second estimate in 2 of Lemma 3.1.

Finally, by $5-7$ of Lemma 3.3 we deduce (22).
Proof of $i$. For $r=+\infty$, let us choose the constants $c^{h}$ in Lemma 3.1 in such a way that $\bar{y}^{h}-\left(x_{1}, h x_{2}, \delta_{h} x_{3}\right)$ has zero average, and let us define $\hat{u}^{h}: \Omega \rightarrow \mathbb{R}^{3}$ by

$$
\begin{equation*}
\hat{u}^{h}:=\frac{\bar{y}^{h}-\left(x_{1}, h x_{2}, \delta_{h} x_{3}\right)}{\varepsilon_{h} / \delta_{h}} \tag{27}
\end{equation*}
$$

From $i$ of Lemma 3.4 we have that $\nabla_{h} \hat{u}^{h}=\frac{\nabla_{h} \bar{y}^{h}-I}{\varepsilon_{h} / \delta_{h}}$ admits a subsequence, not relabeled, converging to $A$ in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$. Hence $\nabla \hat{u}^{h}$ is a Cauchy sequence in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$ and, since $\bar{y}^{h}-\left(x_{1}, h x_{2}, \delta_{h} x_{3}\right)$ has zero average, we have that

$$
\hat{u}^{h} \rightarrow \hat{u} \text { in } W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)
$$

Moreover, since $\nabla_{h} \hat{u}^{h}$ is bounded in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$ and $\nabla_{h} \hat{u}^{h} e_{1}=\hat{u}_{, 1}^{h}$, we deduce that $\hat{u}=\hat{u}\left(x_{1}\right)$ and $A e_{1}=\hat{u}_{, 1}$. In particular, since the matrix $A$ is skewsymmetric, we deduce that $\hat{u}_{1,1}=A_{11}=0, \hat{u}_{2,1}=A_{21}=-A_{12}$ and $\hat{u}_{3,1}=$ $A_{31}=-A_{13}$. By 5 of Lemma 3.3 we have that $A_{12,1}=0$ and hence, $\hat{u}_{2,11}=0$. Putting these information together, and using also the fact that $\hat{u}$ has zero average, we obtain that there exist a constant $\alpha$ and a function $\xi_{3} \in W^{2,2}(0, \ell)$, with $\int_{0}^{\ell} \xi_{3}\left(x_{1}\right) d x_{1}=0$, such that

$$
\begin{equation*}
\hat{u}_{1}=0, \quad \hat{u}_{2}=\alpha\left(x_{1}-\frac{\ell}{2}\right), \quad \hat{u}_{3}=\xi_{3} . \tag{28}
\end{equation*}
$$

On the other hand, using (27) and (13), we have

$$
\int_{\Omega}\left(\hat{u}_{2,1}^{h}-\frac{1}{h} \hat{u}_{1,2}^{h}\right) d x=\frac{1}{\varepsilon_{h} / \delta_{h}} \int_{\Omega}\left(\bar{y}_{2,1}^{h}-\frac{1}{h} \bar{y}_{1,2}^{h}\right) d x=0,
$$

hence, taking the limit as $h \rightarrow 0$, we obtain

$$
0=\int_{\Omega}\left(\hat{u}_{2,1}-A_{12}\right) d x=2 \int_{\Omega} \hat{u}_{2,1} d x
$$

which, in turn, implies that $\alpha=0$ in (28). Hence

$$
\begin{equation*}
\hat{u}_{1}=0, \quad \hat{u}_{2}=0, \quad \hat{u}_{3}=\xi_{3} . \tag{29}
\end{equation*}
$$

Let now $\bar{u}^{h}: \Omega \rightarrow \mathbb{R}^{3}$ be defined by

$$
\bar{y}_{1}^{h}=: x_{1}+\left(\frac{\varepsilon_{h}}{\delta_{h}}\right)^{2} \bar{u}_{1}^{h}, \quad \bar{y}_{2}^{h}=: h x_{2}+\left(\frac{\varepsilon_{h}}{\delta_{h}}\right)^{2} \frac{\bar{u}_{2}^{h}}{h}, \quad \bar{y}_{3}^{h}=: \delta_{h} x_{3}+\left(\frac{\varepsilon_{h}}{\delta_{h}}\right)^{2} \frac{\bar{u}_{3}^{h}}{\delta_{h}}
$$

so that

$$
\frac{\nabla_{h} \bar{y}^{h}-I}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}}=\left(\begin{array}{ccc}
\bar{u}_{1,1}^{h} & \bar{u}_{1,2}^{h} / h & \bar{u}_{1,3}^{h} / \delta_{h}  \tag{30}\\
\bar{u}_{2,1}^{h} / h & \bar{u}_{2,2}^{h} / h^{2} & \bar{u}_{2,3}^{h} /\left(h \delta_{h}\right) \\
\bar{u}_{3,1}^{h} / \delta_{h} & \bar{u}_{3,2}^{h} /\left(h \delta_{h}\right) & \bar{u}_{3,3}^{h} / \delta_{h}^{2}
\end{array}\right)
$$

By (13) we have that

$$
\begin{equation*}
\int_{\Omega}\left(\nabla \bar{u}^{h}-\nabla \bar{u}^{h^{T}}\right) d x=0 . \tag{31}
\end{equation*}
$$

Since we have also that $\int_{\Omega} \bar{u}^{h} d x=0$, by Korn inequality there exists a constant $C_{K}$ such that

$$
\begin{equation*}
\left\|\operatorname{sym} \nabla \bar{u}^{h}\right\|_{L^{2}} \geq C_{K}\left\|\bar{u}^{h}\right\|_{W^{1,2}} \tag{32}
\end{equation*}
$$

By (30) and $i i$ of Lemma 3.4 we have that sym $\nabla \bar{u}^{h}$ admits a Cauchy subsequence in $L^{2}\left(\Omega ; \mathbb{R}^{3}\right)$; hence inequality (32) implies that there exists $\bar{u} \in W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ such that, up to a subsequence,

$$
\bar{u}^{h} \rightarrow \bar{u} \text { in } W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right) .
$$

Moreover, from (30) and $i i$ of Lemma 3.4 it follows that $\bar{u}_{i, \alpha}+\bar{u}_{\alpha, i}=0$ for every $i=1,2,3$ and $\alpha=2,3$, hence $\bar{u}$ is a Bernoulli-Navier displacement. In other words, there exist $\bar{\xi}_{1} \in W^{1,2}(0, \ell)$ and $\bar{\xi}_{2}, \bar{\xi}_{3} \in W^{2,2}(0, \ell)$ such that

$$
\begin{equation*}
\bar{u}_{1}=\bar{\xi}_{1}-x_{2} \bar{\xi}_{2}^{\prime}-x_{3} \bar{\xi}_{3}^{\prime}, \quad \bar{u}_{2}=\bar{\xi}_{2}, \quad \bar{u}_{3}=\bar{\xi}_{3} . \tag{33}
\end{equation*}
$$

Noticing that $\bar{u}_{3}^{h}=\hat{u}_{3}^{h} \delta_{h}^{2} / \varepsilon_{h}$ and recalling that $r=\lim _{h \rightarrow 0} \frac{\varepsilon_{h}}{\delta_{h}^{2}}=+\infty$ yield that $\bar{u}_{3}^{h} \rightarrow 0$ in $W^{1,2}(\Omega)$; hence, $\bar{u}_{3}=\bar{\xi}_{3}=0$. Thus, (33) reduces to

$$
\begin{equation*}
\bar{u}_{1}=\bar{\xi}_{1}-x_{2} \bar{\xi}_{2}^{\prime}, \quad \bar{u}_{2}=\bar{\xi}_{2}, \quad \bar{u}_{3}=0 \tag{34}
\end{equation*}
$$

By (30) and $i i$ of Lemma 3.4 we deduce that $\bar{u}_{1,1}=\frac{1}{2}\left(A^{2}\right)_{11}$. Thus, recalling that $A$ is skew-symmetric, we find

$$
\begin{equation*}
\bar{u}_{1,1}=-\frac{1}{2}\left(\left(A_{21}\right)^{2}+\left(A_{31}\right)^{2}\right)=-\frac{1}{2}\left(\left(\hat{u}_{2,1}\right)^{2}+\left(\hat{u}_{3,1}\right)^{2}\right), \tag{35}
\end{equation*}
$$

and using (29) and (34), we deduce

$$
\bar{\xi}_{1}^{\prime}-x_{2} \bar{\xi}_{2}^{\prime \prime}=-\frac{1}{2}\left(\xi_{3}^{\prime}\right)^{2}
$$

Since the right-hand side depends only on $x_{1}$, this implies

$$
\begin{equation*}
\bar{\xi}_{2}^{\prime \prime}=0 \quad \text { and } \quad \bar{\xi}_{1}^{\prime}=-\frac{1}{2}\left(\xi_{3}^{\prime}\right)^{2} \tag{36}
\end{equation*}
$$

From $\int_{\Omega} \bar{u}_{2} d x=0$ and $\bar{\xi}_{2}^{\prime \prime}=0$, we deduce that $\bar{\xi}_{2}=\bar{k}\left(x_{1}-\frac{\ell}{2}\right)$ for some constant $\bar{k}$. But, as a consequence of (31), we have that

$$
\int_{\Omega}\left(\bar{u}_{1,2}-\bar{u}_{2,1}\right) d x=0
$$

which implies $\bar{k}=0$. Hence, we conclude that

$$
\bar{u}_{1}=\bar{\xi}_{1}, \quad \bar{u}_{2}=0, \quad \bar{u}_{3}=0 .
$$

Moreover, since $\xi_{3} \in W^{2,2}(0, \ell)$, we deduce by (36) that $\bar{\xi}_{1} \in W^{2,2}(0, \ell)$.
The proof of the statement concerning the convergence of $u^{h}$ follows now by the analysis above after setting $u_{1}^{h}:=\bar{u}_{1}^{h}, u_{2}^{h}:=\hat{u}_{2}^{h}, u_{3}^{h}:=\hat{u}_{3}^{h}$, and $\xi_{1}:=\bar{\xi}_{1}$. Finally, since $A_{31}=\hat{u}_{3,1}=\xi_{3}^{\prime}$, equality 6 of Lemma 3.3 implies (24).

Proof of ii. Let now $r \in\{0,1\}$. The proof of this case is very similar to a part of the proof of $i$, thus we only sketch it. Noticing that

$$
\bar{y}_{1}^{h}=x_{1}+\varepsilon_{h} u_{1}^{h}, \quad \bar{y}_{2}^{h}=h x_{2}+\varepsilon_{h} \frac{u_{2}^{h}}{h}, \quad \bar{y}_{3}^{h}=\delta_{h} x_{3}+\varepsilon_{h} \frac{u_{3}^{h}}{\delta_{h}},
$$

we have

$$
\frac{\nabla_{h} \bar{y}^{h}-I}{\varepsilon_{h}}=\left(\begin{array}{ccc}
u_{1,1}^{h} & u_{1,2}^{h} / h & u_{1,3}^{h} / \delta_{h}  \tag{37}\\
u_{2,1}^{h} / h & u_{2,2}^{h} / h^{2} & u_{2,3}^{h} /\left(h \delta_{h}\right) \\
u_{3,1}^{h} / \delta_{h} & u_{3,2}^{h} /\left(h \delta_{h}\right) & u_{3,3}^{h} / \delta_{h}^{2}
\end{array}\right) .
$$

By part iii of Lemma 3.4 and by Korn inequality we deduce that $u^{h} \rightharpoonup u$ in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ with $u \in \mathcal{A}^{B N}$. Moreover, from (37) it also follows that

$$
\frac{\left(\nabla_{h} \bar{y}^{h}\right)_{11}-1}{\varepsilon_{h}} \rightharpoonup u_{1,1} \text { in } L^{2}(\Omega)
$$

Multiplying both sides of (37) by $\delta_{h}$ and using $i$ of Lemma 3.4, we obtain that $A_{12}=0$ and $A_{13}=u_{1,3}$. Passing to the limit in the identity

$$
\left(\bar{R}^{h} G^{h}\right)_{11}=\frac{\left(\nabla_{h} \bar{y}^{h}\right)_{11}-1}{\varepsilon_{h}}-\frac{\bar{R}^{h}{ }_{11}-1}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}} \frac{\varepsilon_{h}}{\delta_{h}^{2}}
$$

after recalling 3 of Lemma 3.1, 2 and 4 of Lemma 3.3, and the definition of $r$, we find

$$
G_{11}=u_{1,1}-r \frac{\left(A^{2}\right)_{11}}{2}=u_{1,1}+r \frac{\left(A_{12}\right)^{2}+\left(A_{13}\right)^{2}}{2}
$$

and this completes the proof.

Remark 3.8 The definitions of $u^{h}$ and $\vartheta^{h}$ in Lemma 3.7 are given in terms of the deformations $\bar{y}^{h}$, which in turn depend on the sequence of constant rotations $\left(Q^{h}\right)$, introduced in Lemma 3.1. By Remark 3.2 any two sequences of constant rotations satisfying Lemma 3.1 have difference going to zero, as $h \rightarrow 0$. Using this fact, one can show that the limits of $\nabla u^{h}$ and $\vartheta^{h}$ are in fact independent of the choice of $\left(Q^{h}\right)$.

Remark 3.9 We give here a geometrical interpretation of $G_{11}$ and of the constraint (21). By explicitly writing $\varepsilon_{h} G_{h}^{T} G_{h} \rightarrow 0$ in $L^{1}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$ we deduce that

$$
\begin{equation*}
\frac{1}{2 \varepsilon_{h}}\left(\nabla_{h} y^{h^{T}} \nabla_{h} y^{h}-I\right) \rightharpoonup \operatorname{sym} G \text { in } L^{1}\left(\Omega ; \mathbb{R}^{3 \times 3}\right) \tag{38}
\end{equation*}
$$

Hence sym $G$ is the limit of a rescaled sequence of Green-St. Venant strain tensors; thus, $G_{11}$ measures the length's variation of fibers parallel to the axis of the beam (see [2]). The component of (38) on the first row and first column can be rewritten as

$$
\begin{equation*}
\frac{\left(\nabla_{h} y^{h}-I\right)_{11}}{\varepsilon_{h}}+\frac{1}{2} \frac{\varepsilon_{h}}{\delta_{h}^{2}}\left[\frac{\left(\nabla_{h} y^{h}-I\right)^{T}}{\varepsilon_{h} / \delta_{h}} \frac{\left(\nabla_{h} y^{h}-I\right)}{\varepsilon_{h} / \delta_{h}}\right]_{11} \rightharpoonup G_{11} \text { in } L^{1}(\Omega) \tag{39}
\end{equation*}
$$

This equation highlights the fact that $G_{11}$ is "generated" by a linear and a quadratic term in $\nabla_{h} y^{h}-I$.

In the case $r \in\{0,1\}$, by using (25) and $i$ of Lemma 3.4, we find

$$
u_{1,1}+\frac{1}{2} r\left(A^{T} A\right)_{11}=G_{11}
$$

that is,

$$
G_{11}(x)=\xi_{1}^{\prime}\left(x_{1}\right)-x_{2} \xi_{2}^{\prime \prime}\left(x_{1}\right)-x_{3} \xi_{3}^{\prime \prime}\left(x_{1}\right)+\frac{r}{2}\left(\xi_{3}^{\prime}\left(x_{1}\right)\right)^{2}
$$

which is exactly (26). We note that when the energy is "small", i.e., $r=0$, the quadratic term in $\nabla_{h} y^{h}-I$ does not give any contribution in $G_{11}$. According to our geometrical interpretation of $G_{11}$, we deduce that the length's variation along the axis of the beam is given by $\xi_{1}^{\prime}$ for $r=0$ and $\xi_{1}^{\prime}+\frac{1}{2}\left(\xi_{3}^{\prime}\right)^{2}$ for $r=1$.

In the case $r=+\infty$, that is, when $\varepsilon_{h} / \delta_{h}^{2} \rightarrow+\infty$, we deduce from (39), after multiplication by $\delta_{h}^{2} / \varepsilon_{h}$, that

$$
\frac{\left(\nabla_{h} y^{h}-I\right)_{11}}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}}+\frac{1}{2}\left[\frac{\left(\nabla_{h} y^{h}-I\right)^{T}}{\varepsilon_{h} / \delta_{h}} \frac{\left(\nabla_{h} y^{h}-I\right)}{\varepsilon_{h} / \delta_{h}}\right]_{11} \rightarrow 0 \text { in } L^{1}(\Omega)
$$

As before, by using (23) and $i$ of Lemma 3.4, we find

$$
u_{1,1}+\frac{1}{2}\left(A^{T} A\right)_{11}=0
$$

that is, $\xi_{1}^{\prime}+\frac{1}{2}\left(\xi_{3}^{\prime}\right)^{2}=0$, which is exactly the constraint (21). This implies that in this regime the axis of the beam is inextensible.

### 3.1 A liminf inequality

In this subsection we prove a lower bound of the limit energy. We start by recalling a result proven in [8, Lemma 3.4].

Lemma 3.10 Assume that $\lim _{h \rightarrow 0} \varepsilon_{h}=0$. Let $\left(y^{h}\right) \subset W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right),\left(R^{h}\right) \subset$ $S O(3)$ and

$$
G^{h}:=\frac{R^{h} \nabla_{h} y^{h}-I}{\varepsilon_{h}} \rightharpoonup G \text { in } L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)
$$

Then

$$
\liminf _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x \geq \frac{1}{2} \int_{\Omega} Q_{3}(G) d x
$$

where $Q_{3}$ is the quadratic form introduced in (3).
We now state and prove the following liminf inequality.
Theorem 3.11 Assume (12). Let $y^{h} \in W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ be a sequence of deformations satisfying (11). Then, there exist rotations $Q^{h} \in S O(3)$ and constants $c^{h} \in \mathbb{R}$ such that, setting $\bar{y}^{h}:=Q^{h^{T}} y^{h}-c^{h}$,

1. $\bar{y}^{h} \rightarrow x_{1} e_{1}$ in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ and $\nabla_{h} \bar{y}^{h} \rightarrow I$ in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$.

Under assumption (18), setting

$$
\begin{aligned}
& u_{1}^{h}:= \begin{cases}\frac{1}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}}\left(\bar{y}_{1}^{h}-x_{1}\right) & \text { if } r=+\infty \\
\frac{1}{\varepsilon_{h}}\left(\bar{y}_{1}^{h}-x_{1}\right) & \text { if } r \in\{0,1\}\end{cases} \\
& u_{2}^{h}:= \begin{cases}\frac{1}{\varepsilon_{h} / \delta_{h}}\left(\bar{y}_{2}^{h}-h x_{2}\right) & \text { if } r=+\infty \\
\frac{1}{\varepsilon_{h} / h}\left(\bar{y}_{2}^{h}-h x_{2}\right) & \text { if } r \in\{0,1\}\end{cases} \\
& u_{3}^{h}:=\frac{1}{\varepsilon_{h} / \delta_{h}}\left(\bar{y}_{3}^{h}-\delta_{h} x_{3}\right), \\
& \vartheta^{h}:=\frac{1}{I_{0}} \frac{1}{h \varepsilon_{h}} \int_{\omega}\left(\delta_{h} x_{2} \bar{y}_{3}^{h}-h x_{3} \bar{y}_{2}^{h}\right) d x_{2} d x_{3},
\end{aligned}
$$

we have that
2. up to subsequences, there exists $u \in \mathcal{A}^{r}$ such that $u^{h} \rightharpoonup u$ in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$; if $r=+\infty$, the convergence is actually strong in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$;
3. up to subsequences, there exists $\vartheta \in W^{1,2}(0, \ell)$ such that $\vartheta^{h} \rightharpoonup \vartheta$ in $W^{1,2}(0, \ell)$.

Moreover,

$$
\begin{equation*}
\liminf _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x \geq I^{r}(u, \vartheta) \tag{40}
\end{equation*}
$$

where $I^{r}: \mathcal{A}^{r} \times W^{1,2}(0, \ell) \rightarrow[0,+\infty)$ is defined by

$$
I^{r}(u, \vartheta):=\frac{1}{24} \int_{0}^{\ell} Q_{2}\left(u_{3}^{\prime \prime}, \vartheta^{\prime}\right) d x_{1}
$$

if $r=+\infty$, and by
$I^{r}(u, \vartheta):=\frac{1}{24} \int_{0}^{\ell} Q_{2}\left(\xi_{3}^{\prime \prime}, \vartheta^{\prime}\right) d x_{1}+\frac{1}{2} \int_{0}^{\ell} E\left(\xi_{1}^{\prime}+\frac{r}{2}\left(\xi_{3}^{\prime}\right)^{2}\right)^{2} d x_{1}+\frac{1}{24} \int_{0}^{\ell} E\left(\xi_{2}^{\prime \prime}\right)^{2} d x_{1}$
if $r \in\{0,1\}$. Here $\xi_{1}, \xi_{2}$, and $\xi_{3}$ are as in the definition of $\mathcal{A}^{r}$ (see Definition 3.6).

Proof. Take as $Q^{h}$ the sequence of rotations constructed in Lemma 3.1 and as $c^{h}$ a sequence of constants chosen as in Lemma 3.7. Then, statement 1 follows from 1 and 3 of Lemma 3.1 and from the fact that $\bar{y}^{h}-\left(x_{1}, h x_{2}, \delta_{h} x_{3}\right)$ has zero average. Statements 2 and 3 follow from Lemma 3.7.

Let us prove (40). Using the frame indifference of $W$ and the definition of $\bar{y}^{h}$ we have that

$$
\begin{equation*}
W\left(\nabla_{h} y^{h}\right)=W\left(\nabla_{h} \bar{y}^{h}\right) . \tag{41}
\end{equation*}
$$

Let $\bar{R}^{h}$ be the sequence of approximating rotations of Lemma 3.1 and let

$$
G^{h}:=\frac{\bar{R}^{h^{T}} \nabla_{h} \bar{y}^{h}-I}{\varepsilon_{h}}
$$

By 4 of Lemma 3.3 we have that, up to subsequences, $G^{h} \rightharpoonup G$ in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$.
Working with the corresponding subsequence (not relabeled) of $\nabla_{h} y^{h}$, and taking into account (41), Lemma 3.10, and (4), we get

$$
\begin{align*}
\liminf _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x & \geq \frac{1}{2} \int_{\Omega} Q_{3}(G) d x=\frac{1}{2} \int_{\Omega} Q_{3}(\operatorname{sym} G) d x \\
& \geq \frac{1}{2} \int_{\Omega} Q_{2}\left(G_{11}, \frac{1}{2}\left(G_{12}+G_{21}\right)\right) d x \tag{42}
\end{align*}
$$

where the last inequality follows from the definition (6) of $Q_{2}$.
By Lemma 3.7 we have that for every $r \in\{0,1,+\infty\}$ there exist $g, \tilde{g} \in$ $L^{2}\left((0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right)\right)$ such that

$$
\begin{gathered}
G_{11}(x)=-x_{3} \xi_{3}^{\prime \prime}\left(x_{1}\right)+g\left(x_{1}, x_{2}\right) \\
\frac{1}{2}\left(G_{12}(x)+G_{21}(x)\right)=-x_{3} \vartheta^{\prime}\left(x_{1}\right)+\tilde{g}\left(x_{1}, x_{2}\right)
\end{gathered}
$$

for a.e. $x \in \Omega$. Since $Q_{2}$ is a quadratic form, we obtain

$$
\begin{align*}
\int_{\Omega} Q_{2}\left(G_{11}, \frac{1}{2}\left(G_{12}+G_{21}\right)\right) d x & =\int_{\Omega} Q_{2}\left(-x_{3} \xi_{3}^{\prime \prime}+g,-x_{3} \vartheta^{\prime}+\tilde{g}\right) d x \\
& =\int_{\Omega} x_{3}^{2} Q_{2}\left(\xi_{3}^{\prime \prime}, \vartheta^{\prime}\right) d x+\int_{\Omega} Q_{2}(g, \tilde{g}) d x \tag{43}
\end{align*}
$$

If $r=+\infty$, we simply deduce

$$
\int_{\Omega} Q_{2}\left(G_{11}, \frac{1}{2}\left(G_{12}+G_{21}\right)\right) d x \geq \int_{\Omega} x_{3}^{2} Q_{2}\left(\xi_{3}^{\prime \prime}, \vartheta^{\prime}\right) d x
$$

hence, by (42)

$$
\liminf _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x \geq \frac{1}{24} \int_{0}^{\ell} Q_{2}\left(\xi_{3}^{\prime \prime}, \vartheta^{\prime}\right) d x_{1}
$$

If $r \in\{0,1\}$, by (26) we have that

$$
g\left(x_{1}, x_{2}\right)=\xi_{1}^{\prime}\left(x_{1}\right)-x_{2} \xi_{2}^{\prime \prime}\left(x_{1}\right)+\frac{r}{2}\left(\xi_{3}^{\prime}\left(x_{1}\right)\right)^{2}
$$

hence, using the definition of $E$ (see (7)) we have

$$
\begin{align*}
\int_{\Omega} Q_{2}(g, \tilde{g}) d x & \geq \int_{\Omega} E\left(\xi_{1}^{\prime}-x_{2} \xi_{2}^{\prime \prime}+\frac{r}{2}\left(\xi_{3}^{\prime}\right)^{2}\right)^{2} d x \\
& =\int_{\Omega} E\left(\xi_{1}^{\prime}+\frac{r}{2}\left(\xi_{3}^{\prime}\right)^{2}\right)^{2} d x+\int_{\Omega} x_{2}^{2} E\left(\xi_{2}^{\prime \prime}\right)^{2} d x \\
& =\int_{0}^{\ell} E\left(\xi_{1}^{\prime}+\frac{r}{2}\left(\xi_{3}^{\prime}\right)^{2}\right)^{2} d x_{1}+\frac{1}{12} \int_{0}^{\ell} E\left(\xi_{2}^{\prime \prime}\right)^{2} d x_{1} \tag{44}
\end{align*}
$$

Thus, combining (42) - (44), we conclude that

$$
\begin{aligned}
& \liminf _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x \geq \\
& \quad \geq \frac{1}{24} \int_{0}^{\ell} Q_{2}\left(\xi_{3}^{\prime \prime}, \vartheta^{\prime}\right) d x_{1}+\frac{1}{2} \int_{0}^{\ell} E\left(\xi_{1}^{\prime}+\frac{r}{2}\left(\xi_{3}^{\prime}\right)^{2}\right)^{2} d x_{1}+\frac{1}{24} \int_{0}^{\ell} E\left(\xi_{2}^{\prime \prime}\right)^{2} d x_{1}
\end{aligned}
$$

### 3.2 Recovery sequences

Here we shall prove that the lower bound obtained in the previous subsection is achieved. For clarity we shall discuss the cases $r=\infty$ and $r \in\{0,1\}$ in two different subsections.

### 3.2.1 The recovery sequence in the case $r=\infty$

In this subsection we consider the case in which

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{\varepsilon_{h}}{\delta_{h}}=0 \quad \text { and } \quad \lim _{h \rightarrow 0} \frac{\varepsilon_{h}}{\delta_{h}^{2}}=r=+\infty \tag{45}
\end{equation*}
$$

and we further assume that

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{h^{2} \varepsilon_{h}}{\delta_{h}^{2}}=0 \tag{46}
\end{equation*}
$$

Theorem 3.12 Assume (45) and (46). Then for every $(u, \vartheta) \in \mathcal{A}^{\infty} \times W^{1,2}(0, \ell)$ there exists a sequence of deformations $y^{h} \in W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ such that, setting

$$
\begin{aligned}
u_{1}^{h} & :=\frac{1}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}}\left(y_{1}^{h}-x_{1}\right) \\
u_{2}^{h} & :=\frac{1}{\varepsilon_{h} / \delta_{h}}\left(y_{2}^{h}-h x_{2}\right) \\
u_{3}^{h} & :=\frac{1}{\varepsilon_{h} / \delta_{h}}\left(y_{3}^{h}-\delta_{h} x_{3}\right) \\
\vartheta^{h} & :=\frac{1}{I_{0}} \frac{1}{h \varepsilon_{h}} \int_{\omega}\left(\delta_{h} x_{2} y_{3}^{h}-h x_{3} y_{2}^{h}\right) d x_{2} d x_{3}
\end{aligned}
$$

we have that $\nabla_{h} y^{h} \rightarrow I$ in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right), u^{h} \rightarrow u$ in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$, $\vartheta^{h} \rightarrow \vartheta$ in $W^{1,2}(0, \ell)$, and

$$
\begin{equation*}
\limsup _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x \leq I^{\infty}(u, \vartheta) \tag{47}
\end{equation*}
$$

where

$$
I^{\infty}(u, \vartheta):=\frac{1}{24} \int_{0}^{\ell} Q_{2}\left(u_{3}^{\prime \prime}, \vartheta^{\prime}\right) d x_{1}
$$

Proof. Let us fix $(u, \vartheta) \in \mathcal{A}^{\infty} \times W^{1,2}(0, \ell)$ smooth enough and let $\xi_{1}$ and $\xi_{3}$ be as in the definition of $\mathcal{A}^{\infty}$, see Definition 3.5. For every $t \in[0, \ell]$ we define

$$
A(t):=\left(\begin{array}{ccc}
0 & 0 & -\xi_{3}^{\prime}(t) \\
0 & 0 & -\vartheta(t) \\
\xi_{3}^{\prime}(t) & \vartheta(t) & 0
\end{array}\right)
$$

To simplify notation we set $\eta_{h}:=\varepsilon_{h} / \delta_{h}$, which tends to 0 by (45). Let $R^{h}:[0, \ell] \rightarrow \mathbb{R}^{3 \times 3}$ be the solution of the Cauchy problem

$$
\begin{cases}X^{\prime}=\eta_{h} X A^{\prime} & \text { in }[0, \ell] \\ X(0)=\exp \left(\eta_{h} A(0)\right) & \end{cases}
$$

It is easy to see that $R^{h}(t) \in S O(3)$ for every $t$; indeed, $R^{h}(0) \in S O(3)$ and $\left(R^{h}\left(R^{h}\right)^{T}\right)^{\prime}=0$ on $[0, \ell]$ from the equation. Moreover, the function

$$
Q^{h}(t):=I+\eta_{h} A(t)+\eta_{h}^{2} \int_{0}^{t} A(s) A^{\prime}(s) d s+\frac{1}{2} \eta_{h}^{2} A^{2}(0)
$$

solves the problem

$$
\begin{cases}\left(Q^{h}\right)^{\prime}=\eta_{h} Q^{h} A^{\prime}+\eta_{h}^{3} L & \text { in }[0, \ell] \\ Q^{h}(0)=\exp \left(\eta_{h} A(0)\right)+O\left(\eta_{h}^{3}\right), & \end{cases}
$$

where

$$
L(t)=-\left(\int_{0}^{t} A(s) A^{\prime}(s) d s+\frac{1}{2} A^{2}(0)\right) A^{\prime}(t)
$$

Therefore, by Gronwall Lemma we have that $\left|R^{h}-Q^{h}\right|=O\left(\eta_{h}^{3}\right)$ uniformly on $[0, \ell]$; in other words,

$$
\begin{equation*}
R^{h}(t)=I+\eta_{h} A(t)+\eta_{h}^{2} \int_{0}^{t} A(s) A^{\prime}(s) d s+\frac{1}{2} \eta_{h}^{2} A^{2}(0)+O\left(\eta_{h}^{3}\right) \tag{48}
\end{equation*}
$$

Finally, let us fix $\varphi \in C^{\infty}([0, \ell])$ and $\gamma \in C^{\infty}\left([0, \ell] ; \mathbb{R}^{3}\right)$, and define

$$
\begin{aligned}
\beta^{h}(x):= & \varepsilon_{h} R^{h}\left(x_{1}\right)\left(\begin{array}{c}
-\frac{1}{2} h^{2} x_{2}^{2} x_{3} \varphi^{\prime}\left(x_{1}\right) \\
-h x_{2} x_{3} \varphi\left(x_{1}\right) \\
\frac{1}{2} \frac{h^{2}}{\delta_{h}} x_{2}^{2} \varphi\left(x_{1}\right)
\end{array}\right) \\
& -\varepsilon_{h} h x_{2} x_{3} \vartheta^{\prime}\left(x_{1}\right) R^{h}\left(x_{1}\right) e_{1}+\frac{1}{2} \varepsilon_{h} \delta_{h} x_{3}^{2} R^{h}\left(x_{1}\right) \gamma\left(x_{1}\right) .
\end{aligned}
$$

We consider the sequence of three-dimensional deformations $y^{h}: \Omega \rightarrow \mathbb{R}^{3}$ given by

$$
y^{h}(x):=\int_{0}^{x_{1}} R^{h}(t) e_{1} d t+R^{h}\left(x_{1}\right)\left(\begin{array}{c}
0 \\
h x_{2} \\
\delta_{h} x_{3}
\end{array}\right)+\beta^{h}(x)+c^{h}
$$

where $c_{1}^{h}:=\eta_{h}^{2} \xi_{1}(0), c_{2}^{h}:=0$ and $c_{3}^{h}:=\eta_{h} \xi_{3}(0)$.
We first check the convergence of the tangential and normal displacements $u^{h}$. The expansion (48) and the definition of $\mathcal{A}^{\infty}$ imply that

$$
\begin{align*}
R_{11}^{h}\left(x_{1}\right) & =1-\eta_{h}^{2} \int_{0}^{x_{1}} \xi_{3}^{\prime}(s) \xi_{3}^{\prime \prime}(s) d s-\frac{1}{2} \eta_{h}^{2} \xi_{3}^{\prime}(0)^{2}+O\left(\eta_{h}^{3}\right) \\
& =1-\frac{1}{2} \eta_{h}^{2} \xi_{3}^{\prime}\left(x_{1}\right)^{2}+O\left(\eta_{h}^{3}\right) \\
& =1+\eta_{h}^{2} \xi_{1}^{\prime}\left(x_{1}\right)+O\left(\eta_{h}^{3}\right) \tag{49}
\end{align*}
$$

and

$$
\begin{equation*}
R_{12}^{h}\left(x_{1}\right)=O\left(\eta_{h}^{2}\right), \quad R_{13}^{h}\left(x_{1}\right)=O\left(\eta_{h}\right), \tag{50}
\end{equation*}
$$

as $h \rightarrow 0$, uniformly in $[0, \ell]$. Since $R^{h}=I+O\left(\eta_{h}\right)$, we have that $\beta_{1}^{h}=o\left(\eta_{h}^{2}\right)$ and $\nabla \beta_{1}^{h}=o\left(\eta_{h}^{2}\right)$. Combining these two facts with (48)-(50), we obtain that

$$
\begin{aligned}
u_{1}^{h}(x) & =\xi_{1}\left(x_{1}\right)+o(1) \\
\nabla u_{1}^{h}(x) & =\xi_{1}^{\prime}\left(x_{1}\right) e_{1}+o(1)
\end{aligned}
$$

Therefore, we can conclude that $u_{1}^{h}$ converges to $u_{1}$ strongly in $W^{1,2}(\Omega)$.
Similar computations show that $u_{k}^{h}$ converges to $u_{k}$ strongly in $W^{1,2}(\Omega)$ for $k=2,3$.

Finally, we note that

$$
\vartheta^{h}=\vartheta+\frac{1}{I_{0}} \frac{1}{h \varepsilon_{h}} \int_{\omega}\left(\delta_{h} x_{2} \beta_{3}^{h}-h x_{3} \beta_{2}^{h}\right) d x_{2} d x_{3}+O\left(\eta_{h}\right)=\vartheta+o(1)
$$

which gives the desired convergence.
Let us prove now the convergence of energies (47). By differentiation we obtain

$$
\nabla_{h} y^{h}=R^{h}\left(x_{1}\right)+\left(R^{h}\right)^{\prime}\left(x_{1}\right)\left(\begin{array}{c}
0 \\
h x_{2} \\
\delta_{h} x_{3}
\end{array}\right) \otimes e_{1}+\nabla_{h} \beta^{h}
$$

Since by definition

$$
\begin{equation*}
R^{h^{T}}\left(R^{h}\right)^{\prime}=\eta_{h} A^{\prime} \tag{51}
\end{equation*}
$$

and $\eta_{h} \delta_{h}=\varepsilon_{h}$, we deduce that

$$
R^{h^{T}} \nabla_{h} y^{h}=I+\varepsilon_{h}\left(\begin{array}{c}
-x_{3} \xi_{3}^{\prime \prime} \\
-x_{3} \vartheta^{\prime} \\
\frac{h}{\delta_{h}} x_{2} \vartheta^{\prime}
\end{array}\right) \otimes e_{1}+R^{h^{T}} \nabla_{h} \beta^{h}
$$

Using property (51) and the orthogonality of $R^{h}$, a direct computation shows that

$$
\begin{align*}
R^{h^{T}} \nabla_{h} \beta^{h}= & \varepsilon_{h}\left(\begin{array}{ccc}
0 & -x_{3} \vartheta^{\prime} & -\frac{1}{2} \frac{h^{2}}{\delta_{h}} x_{2}^{2} \varphi^{\prime}-\frac{h}{\delta_{h}} x_{2} \vartheta^{\prime} \\
0 & -x_{3} \varphi & -\frac{h}{\delta_{h}} x_{2} \varphi \\
\frac{1}{2} \frac{h^{2}}{\delta_{h}} x_{2}^{2} \varphi^{\prime} & \frac{h}{\delta_{h}} x_{2} \varphi & 0
\end{array}\right)  \tag{52}\\
& +\varepsilon_{h} x_{3}\left(\gamma \otimes e_{3}\right)+\frac{1}{2} \frac{h^{2} \varepsilon_{h}^{2}}{\delta_{h}^{2}} x_{2}^{2} \varphi A^{\prime} e_{3} \otimes e_{1}+o\left(\varepsilon_{h}\right)
\end{align*}
$$

Under the assumption (46) the second term in the second line of the previous formula is of order $o\left(\varepsilon_{h}\right)$. Hence

$$
R^{h^{T}} \nabla_{h} y^{h}=I+\varepsilon_{h} B^{h}(x)+o\left(\varepsilon_{h}\right),
$$

where

$$
B^{h}:=\left(\begin{array}{ccc}
-x_{3} \xi_{3}^{\prime \prime} & -x_{3} \vartheta^{\prime} & -\frac{1}{2} \frac{h^{2}}{\delta_{h}} x_{2}^{2} \varphi^{\prime}-\frac{h}{\delta_{h}} x_{2} \vartheta^{\prime}+x_{3} \gamma_{1} \\
-x_{3} \vartheta^{\prime} & -x_{3} \varphi & -\frac{h}{\delta_{h}} x_{2} \varphi+x_{3} \gamma_{2} \\
\frac{h}{\delta_{h}} x_{2} \vartheta^{\prime}+\frac{1}{2} \frac{h^{2}}{\delta_{h}} x_{2}^{2} \varphi^{\prime} & \frac{h}{\delta_{h}} x_{2} \varphi & x_{3} \gamma_{3}
\end{array}\right)
$$

Applying the identity $(I+B)^{T}(I+B)=I+2 \operatorname{sym} B+B^{T} B$ and observing that $\varepsilon_{h}^{2}\left(B^{h}\right)^{T} B^{h}=O\left(\varepsilon_{h}^{2} h^{2} / \delta_{h}^{2}\right)=o\left(\varepsilon_{h}\right)$ by (46), we obtain

$$
\left(\nabla_{h} y^{h}\right)^{T} \nabla_{h} y^{h}=I+2 \varepsilon_{h} x_{3} Z+o\left(\varepsilon_{h}\right)
$$

where

$$
Z=\left(\begin{array}{ccc}
-\xi_{3}^{\prime \prime} & -\vartheta^{\prime} & \frac{1}{2} \gamma_{1} \\
-\vartheta^{\prime} & -\varphi & \frac{1}{2} \gamma_{2} \\
\frac{1}{2} \gamma_{1} & \frac{1}{2} \gamma_{2} & \gamma_{3}
\end{array}\right)
$$

By frame-indifference we have

$$
W\left(\nabla_{h} y^{h}\right)=W\left(\sqrt{\left(\nabla_{h} y^{h}\right)^{T} \nabla_{h} y^{h}}\right)=W\left(I+\varepsilon_{h} x_{3} Z+o\left(\varepsilon_{h}\right)\right)
$$

As $Z$ is bounded in $L^{\infty}$, for $h$ small enough the matrix $I+\varepsilon_{h} x_{3} Z+o\left(\varepsilon_{h}\right)$ belongs to the neighborhood of $S O(3)$ where $W$ is of class $C^{2}$, so that, by expanding $W$ around the identity, we have

$$
\begin{gathered}
\varepsilon_{h}^{-2} W\left(\nabla_{h} y^{h}\right) \rightarrow \frac{1}{2} x_{3}^{2} Q_{3}(Z) \text { a.e. in } \Omega, \\
\varepsilon_{h}^{-2}\left|W\left(\nabla_{h} y^{h}\right)\right| \leq C\left(|Z|^{2}+1\right)
\end{gathered}
$$

By the dominated convergence theorem this implies

$$
\begin{equation*}
\lim _{h} \frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x=\frac{1}{24} \int_{0}^{\ell} Q_{3}(Z) d x_{1} . \tag{53}
\end{equation*}
$$

Consider now the general case. Let $(u, \vartheta) \in \mathcal{A}^{\infty} \times W^{1,2}(0, \ell)$, and let $\xi_{1}$ and $\xi_{3}$ be as in the definition of $\mathcal{A}^{\infty}$. Let also $\varphi \in L^{2}(0, \ell)$ and $\gamma \in L^{2}\left((0, \ell) ; \mathbb{R}^{3}\right)$ be such that

$$
Q_{2}\left(u_{3}^{\prime \prime}, \vartheta^{\prime}\right)=Q_{3}\left(\begin{array}{ccc}
-\xi_{3}^{\prime \prime} & -\vartheta^{\prime} & \frac{1}{2} \gamma_{1}  \tag{54}\\
-\vartheta^{\prime} & -\varphi & \frac{1}{2} \gamma_{2} \\
\frac{1}{2} \gamma_{1} & \frac{1}{2} \gamma_{2} & \gamma_{3}
\end{array}\right)
$$

a.e. in $(0, \ell)$. Finally, let us consider sequences $\xi_{3}^{k}, \vartheta^{k}, \varphi^{k} \in C^{\infty}([0, \ell])$ and $\gamma^{k} \in C^{\infty}\left([0, \ell] ; \mathbb{R}^{3}\right)$ such that $\xi_{3}^{k} \rightarrow \xi_{3}$ strongly in $W^{2,2}(0, \ell), \vartheta^{k} \rightarrow \vartheta$ strongly in
$W^{1,2}(0, \ell), \varphi^{k} \rightarrow \varphi$ strongly in $L^{2}(0, \ell)$, and $\gamma^{k} \rightarrow \gamma$ strongly in $L^{2}\left((0, \ell) ; \mathbb{R}^{3}\right)$. Setting $u^{k}:=\left(\xi_{1}^{k}, 0, \xi_{3}^{k}\right)$, where

$$
\xi_{1}^{k}\left(x_{1}\right):=-\frac{1}{2} \int_{0}^{x_{1}}\left(\left(\xi_{3}^{k}\right)^{\prime}(t)\right)^{2} d t+\xi_{1}(0)
$$

for every $x_{1} \in(0, \ell)$, it is immediate to see that $u^{k} \in \mathcal{A}^{\infty} \cap C^{\infty}\left([0, \ell] ; \mathbb{R}^{3}\right)$ and $u^{k} \rightarrow u$ in $W^{2,2}\left((0, \ell) ; \mathbb{R}^{3}\right)$. By the previous argument for every $k \in \mathbb{N}$ we can construct a sequence of three-dimensional deformations, whose associated displacement and twist function converge to $\left(u^{k}, \vartheta^{k}\right)$, as $h \rightarrow 0$, and satisfying (53) with $Z$ replaced by

$$
Z^{k}:=\left(\begin{array}{ccc}
-\left(\xi_{3}^{k}\right)^{\prime \prime} & -\left(\vartheta^{k}\right)^{\prime} & \frac{1}{2} \gamma_{1}^{k} \\
-\left(\vartheta^{k}\right)^{\prime} & -\varphi^{k} & \frac{1}{2} \gamma_{2}^{k} \\
\frac{1}{2} \gamma_{1}^{k} & \frac{1}{2} \gamma_{2}^{k} & \gamma_{3}^{k}
\end{array}\right)
$$

Using a diagonal argument, the continuity of the left-handside of (53) with respect to the $L^{2}$ convergence, and equality (54), we deduce the $\Gamma$-limsup inequality (47).

Remark 3.13 The assumption (46) used in Theorem 3.12 is crucial in the construction of the recovery sequence since it allows us to control the stretch of the mid-plane, i.e., the $x_{1} x_{2}$-plane. For instance, in (52) it permits to drop the term

$$
\frac{1}{2} \frac{h^{2} \varepsilon_{h}^{2}}{\delta_{h}^{2}} x_{2}^{2} \varphi A^{\prime} e_{3} \otimes e_{1}=\frac{1}{2} \frac{h^{2} \varepsilon_{h}^{2}}{\delta_{h}^{2}} x_{2}^{2} \varphi\left(-\xi_{3}^{\prime \prime} e_{1} \otimes e_{1}-\vartheta^{\prime} e_{2} \otimes e_{1}\right)
$$

that clearly represents a mid-plane deformation.
We note also that (46) coincides with the assumption that was required in the analysis of the critical regime, developed in [7]. Indeed, in this case we have $\lim _{h \rightarrow 0} \frac{\varepsilon_{h}}{\delta_{h}}=1$, so that (46) is equivalent to

$$
\lim _{h \rightarrow 0} \frac{h^{2}}{\delta_{h}}=0
$$

which coincides with condition (5.6) in [7].

### 3.2.2 The recovery sequence in the case $r \in\{0,1\}$

In this subsection we consider the case in which

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{\varepsilon_{h}}{\delta_{h}}=0 \quad \text { and } \quad \lim _{h \rightarrow 0} \frac{\varepsilon_{h}}{\delta_{h}^{2}}=r \in\{0,1\} \tag{55}
\end{equation*}
$$

Theorem 3.14 Assume (55). Then for every $(u, \vartheta) \in \mathcal{A}^{B N} \times W^{1,2}(0, \ell)$ there exists a sequence of deformations $y^{h} \in W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ such that, setting

$$
\begin{aligned}
u_{1}^{h} & :=\frac{1}{\varepsilon_{h}}\left(y_{1}^{h}-x_{1}\right) \\
u_{2}^{h} & :=\frac{1}{\varepsilon_{h} / h}\left(y_{2}^{h}-h x_{2}\right) \\
u_{3}^{h} & :=\frac{1}{\varepsilon_{h} / \delta_{h}}\left(y_{3}^{h}-\delta_{h} x_{3}\right) \\
\vartheta^{h} & :=\frac{1}{I_{0}} \frac{1}{h \varepsilon_{h}} \int_{\omega}\left(\delta_{h} x_{2} y_{3}^{h}-h x_{3} y_{2}^{h}\right) d x_{2} d x_{3}
\end{aligned}
$$

we have that $\nabla_{h} y^{h} \rightarrow I$ in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right), u^{h} \rightarrow u$ in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right), \vartheta^{h} \rightarrow \vartheta$ in $W^{1,2}(0, \ell)$, and

$$
\begin{equation*}
\limsup _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x \leq I^{r}(u, \vartheta) \tag{56}
\end{equation*}
$$

where
$I^{r}(u, \vartheta):=\frac{1}{24} \int_{0}^{\ell} Q_{2}\left(\xi_{3}^{\prime \prime}, \vartheta^{\prime}\right) d x_{1}+\frac{1}{2} \int_{0}^{\ell} E\left(\xi_{1}^{\prime}+\frac{r}{2}\left(\xi_{3}^{\prime}\right)^{2}\right)^{2} d x_{1}+\frac{1}{24} \int_{0}^{\ell} E\left(\xi_{2}^{\prime \prime}\right)^{2} d x_{1}$ with $\xi_{1}, \xi_{2}$ and $\xi_{3}$ as in the definition of $\mathcal{A}^{B N}$.

Proof. Let us fix $(u, \vartheta) \in \mathcal{A}^{B N} \times W^{1,2}(0, \ell)$ smooth enough and let $\xi_{1}, \xi_{2}$, and $\xi_{3}$ be as in the definition of $\mathcal{A}^{B N}$. Let us fix $\alpha, \gamma, \sigma \in C^{\infty}\left([0, \ell] ; \mathbb{R}_{s y m}^{3 \times 3}\right)$.

We consider the sequence of three-dimensional deformations $y^{h}: \Omega \rightarrow \mathbb{R}^{3}$ given by

$$
y^{h}(x):=\left(\begin{array}{c}
x_{1} \\
h x_{2} \\
\delta_{h} x_{3}
\end{array}\right)+\varepsilon_{h}\left(\begin{array}{c}
\xi_{1}-x_{2} \xi_{2}^{\prime}-x_{3} \xi_{3}^{\prime}-h x_{2} x_{3} \vartheta^{\prime} \\
\frac{1}{h} \xi_{2}-x_{3} \vartheta \\
\frac{1}{\delta_{h}} \xi_{3}+\frac{h}{\delta_{h}} x_{2} \vartheta
\end{array}\right)+\varepsilon_{h} \beta^{h}(x)
$$

where

$$
\left\{\begin{array}{l}
\beta_{1}^{h}:=\delta_{h}\left(x_{3}^{2} \alpha_{13}+2 x_{2} x_{3} \gamma_{13}+2 x_{3} \sigma_{13}\right)+\frac{1}{2} h^{2} x_{2}^{2} x_{3} \alpha_{22}^{\prime}+h\left(x_{2}^{2} \gamma_{12}+2 x_{2} \sigma_{12}\right) \\
\beta_{2}^{h}:=h\left(x_{2} x_{3} \alpha_{22}+\frac{1}{2} x_{2}^{2} \gamma_{22}+x_{2} \sigma_{22}\right)+\delta_{h}\left(x_{3}^{2} \alpha_{23}+2 x_{3} \sigma_{23}\right) \\
\beta_{3}^{h}:=\delta_{h}\left(\frac{1}{2} x_{3}^{2} \alpha_{33}+x_{2} x_{3} \gamma_{33}+x_{3} \sigma_{33}\right)+h x_{2}^{2} \gamma_{23}-\frac{1}{2} \frac{h^{2}}{\delta_{h}} x_{2}^{2} \alpha_{22}
\end{array}\right.
$$

It is easy to see that the displacement and the twist function associated with $y^{h}$ converge to $u$ and $\vartheta$ in $W^{1,2}$. Moreover, we have that

$$
\begin{equation*}
\nabla_{h} y^{h}=I+\varepsilon_{h}\left(M^{h}+\frac{1}{\delta_{h}} A+\nabla_{h} \beta^{h}\right) \tag{57}
\end{equation*}
$$

where

$$
M^{h}:=\left(\begin{array}{ccc}
u_{1,1}-h x_{2} x_{3} \vartheta^{\prime \prime} & -\xi_{2}^{\prime} / h-x_{3} \vartheta^{\prime} & -h x_{2} \vartheta^{\prime} / \delta_{h} \\
\xi_{2}^{\prime} / h-x_{3} \vartheta^{\prime} & 0 & 0 \\
h x_{2} \vartheta^{\prime} / \delta_{h} & 0 & 0
\end{array}\right)
$$

and

$$
A:=\left(\begin{array}{ccc}
0 & 0 & -\xi_{3}^{\prime} \\
0 & 0 & -\vartheta \\
\xi_{3}^{\prime} & \vartheta & 0
\end{array}\right)
$$

Also,

$$
\begin{aligned}
\nabla_{h} \beta^{h}= & \left(\begin{array}{ccc}
0 & x_{2} \gamma_{12}+\sigma_{12} & x_{3} \alpha_{13}+x_{2} \gamma_{13}+\sigma_{13} \\
& x_{3} \alpha_{22}+x_{2} \gamma_{22}+\sigma_{22} & x_{3} \alpha_{23}+x_{2} \gamma_{23}+\sigma_{23} \\
\operatorname{sym} & x_{3} \alpha_{33}+x_{2} \gamma_{33}+\sigma_{33}
\end{array}\right) \\
& +\left(\begin{array}{ccc}
0 & x_{2} \gamma_{12}+\sigma_{12} & x_{3} \alpha_{13}+x_{2} \gamma_{13}+\sigma_{13}+h^{2} x_{2}^{2} \alpha_{22}^{\prime} /\left(2 \delta_{h}\right) \\
\text { skw } & 0 & h x_{2} \alpha_{22} / \delta_{h}+x_{3} \alpha_{23}-x_{2} \gamma_{23}+\sigma_{23} \\
\text { skw }
\end{array}\right) \\
& +O(h)+O\left(\delta_{h} / h\right) .
\end{aligned}
$$

Thus, we obtain

$$
\begin{aligned}
\nabla_{h} y^{h^{T}} \nabla_{h} y^{h}= & I+2 \varepsilon_{h} \operatorname{sym}\left(M^{h}+\frac{1}{\delta_{h}} A+\nabla_{h} \beta^{h}\right)+\frac{\varepsilon_{h}^{2}}{\delta_{h}^{2}} A^{T} A \\
& +O\left(\frac{\varepsilon_{h}^{2}}{h \delta_{h}}\right)+O\left(\frac{h \varepsilon_{h}^{2}}{\delta_{h}^{2}}\right) \\
= & I+2 \varepsilon_{h} \operatorname{sym}\left(M^{h}+\nabla_{h} \beta^{h}+\frac{r}{2} A^{T} A\right)+o\left(\varepsilon_{h}\right)
\end{aligned}
$$

where the last equality follows from (55). From the above relations we deduce that

$$
\operatorname{sym}\left(M^{h}+\nabla_{h} \beta^{h}+\frac{r}{2} A^{T} A\right)=Z_{1}+x_{2} Z_{2}+x_{3} Z_{3}+O(h)+O\left(\delta_{h} / h\right)
$$

where

$$
\begin{gather*}
Z_{1}:=\left(\begin{array}{ccc}
\xi_{1}^{\prime}+\frac{r}{2}\left(\xi_{3}^{\prime}\right)^{2} & \frac{r}{2} \xi_{3}^{\prime} \vartheta+\sigma_{12} & \sigma_{13} \\
& \frac{r}{2} \vartheta^{2}+\sigma_{22} & \sigma_{23} \\
\operatorname{sym} & \frac{r}{2}\left(\left(\xi_{3}^{\prime}\right)^{2}+\vartheta^{2}\right)+\sigma_{33}
\end{array}\right)  \tag{58}\\
Z_{2}:=\left(\begin{array}{ccc}
-\xi_{2}^{\prime \prime} & \gamma_{12} & \gamma_{13} \\
& \gamma_{22} & \gamma_{23} \\
\operatorname{sym} & & \gamma_{33}
\end{array}\right), \quad Z_{3}:=\left(\begin{array}{ccc}
-\xi_{3}^{\prime \prime} & -\vartheta^{\prime} & \alpha_{13} \\
& \alpha_{22} & \alpha_{23} \\
\operatorname{sym} & & \alpha_{33}
\end{array}\right) . \tag{59}
\end{gather*}
$$

By frame-indifference we have

$$
W\left(\nabla_{h} y^{h}\right)=W\left(\sqrt{\left(\nabla_{h} y^{h}\right)^{T} \nabla_{h} y^{h}}\right)=W\left(I+\varepsilon_{h}\left(Z_{1}+x_{2} Z_{2}+x_{3} Z_{3}\right)+o\left(\varepsilon_{h}\right)\right)
$$

As $\left(Z_{1}+x_{2} Z_{2}+x_{3} Z_{3}\right) \in L^{\infty}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$, for $h$ small enough the matrix $I+$ $\varepsilon_{h}\left(Z_{1}+x_{2} Z_{2}+x_{3} Z_{3}\right)+o\left(\varepsilon_{h}\right)$ belongs to the neighborhood of $S O(3)$ where $W$ is of class $C^{2}$, so that by Taylor expansion we have

$$
\begin{gathered}
\varepsilon_{h}^{-2} W\left(\nabla_{h} y^{h}\right) \rightarrow \frac{1}{2} Q_{3}\left(Z_{1}+x_{2} Z_{2}+x_{3} Z_{3}\right) \text { a.e. in } \Omega \\
\quad \varepsilon_{h}^{-2}\left|W\left(\nabla_{h} y^{h}\right)\right| \leq C\left(\left|Z_{1}+x_{2} Z_{2}+x_{3} Z_{3}\right|^{2}+1\right)
\end{gathered}
$$

By the dominated convergence theorem this implies

$$
\begin{aligned}
& \lim _{h} \frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x= \\
& \quad=\frac{1}{2} \int_{\Omega} Q_{3}\left(Z_{1}+x_{2} Z_{2}+x_{3} Z_{3}\right) d x \\
& \quad=\frac{1}{2} \int_{\Omega} Q_{3}\left(Z_{1}\right)+x_{2}^{2} Q_{3}\left(Z_{2}\right)+x_{3}^{2} Q_{3}\left(Z_{3}\right) d x \\
& \quad=\frac{1}{24} \int_{0}^{\ell} Q_{3}\left(Z_{3}\right) d x_{1}+\frac{1}{2} \int_{0}^{\ell} Q_{3}\left(Z_{1}\right) d x_{1}+\frac{1}{24} \int_{0}^{\ell} Q_{3}\left(Z_{2}\right) d x_{1}
\end{aligned}
$$

Consider now the general case. Let $(u, \vartheta) \in \mathcal{A}^{B N} \times W^{1,2}(0, \ell)$, and let $\xi_{1}$, $\xi_{2}$, and $\xi_{3}$ be as in the definition of $\mathcal{A}^{B N}$. Let also $\alpha, \gamma, \sigma \in L^{2}\left((0, \ell) ; \mathbb{R}_{s y m}^{3 \times 3}\right)$ be such that

$$
E\left(\xi_{1}^{\prime}+\frac{r}{2}\left(\xi_{3}^{\prime}\right)^{2}\right)^{2}=Q_{3}\left(Z_{1}\right), \quad E\left(\xi_{2}^{\prime \prime}\right)^{2}=Q_{3}\left(Z_{2}\right), \quad Q_{2}\left(\xi_{3}^{\prime \prime}, \vartheta^{\prime}\right)=Q_{3}\left(Z_{3}\right)
$$

a.e. in $(0, \ell)$, where the $Z_{i}$ are defined as in (58)-(59). Arguing as in the proof of Theorem 3.12, we deduce the $\Gamma$-limsup inequality (56) by density.

## 4 Convergence of minimizers

In this section we introduce a sequence of forces and characterize the asymptotic behaviour, as $h \rightarrow 0$, of minimizers (or almost minimizers) of the total energy. This is made precise in the following theorem.

Theorem 4.1 Let $\left(\alpha_{h}\right)$ be a sequence of strictly positive real numbers such that

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{\alpha_{h}}{\delta_{h}^{2}}=0 \tag{60}
\end{equation*}
$$

and let $\varepsilon_{h}:=\alpha_{h} / \delta_{h}$. Let $f^{h}, f \in L^{2}\left((0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right)\right)$ be such that

$$
\begin{equation*}
\int_{(0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right)} f^{h} d x_{1} d x_{2}=\int_{(0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right)} x_{i} f^{h} d x_{1} d x_{2}=0, \quad i=1,2 \tag{61}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{1}{\alpha_{h}} f^{h} \rightharpoonup f \text { weakly in } L^{2}\left((0, \ell) \times\left(-\frac{1}{2}, \frac{1}{2}\right)\right) \tag{62}
\end{equation*}
$$

Let

$$
J^{h}(y):=I^{h}(y)-\int_{\Omega} f^{h} e_{3} \cdot y d x
$$

for $y \in W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$.
Then the following statements hold:
i. $\left|\inf J^{h}\right| \leq C \varepsilon_{h}^{2}$ for any $h>0$ small enough.
ii. If $\left(y^{h}\right) \subset W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ is a minimizing sequence of $\frac{1}{\varepsilon_{h}^{2}} J^{h}$ in the following sense

$$
\begin{equation*}
\lim _{h \rightarrow 0}\left(\frac{1}{\varepsilon_{h}^{2}} J^{h}\left(y^{h}\right)-\inf \frac{1}{\varepsilon_{h}^{2}} J^{h}\right)=0 \tag{63}
\end{equation*}
$$

then there exist some constants $Q^{h} \in S O(3)$ and $c^{h} \in \mathbb{R}^{3}$ such that, setting $\bar{y}^{h}:=Q^{h^{T}} y^{h}-c^{h}$, we have that

$$
\begin{equation*}
\bar{y}^{h} \rightarrow x_{1} e_{1} \quad \text { strongly in } W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right) . \tag{64}
\end{equation*}
$$

iii. Assume in addition that

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{\alpha_{h}}{\delta_{h}^{3}}=\lim _{h \rightarrow 0} \frac{\varepsilon_{h}}{\delta_{h}^{2}}=r \in\{0,1,+\infty\} . \tag{65}
\end{equation*}
$$

If $r=+\infty$, assume also that

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{h^{2} \varepsilon_{h}}{\delta_{h}^{2}}=0 . \tag{66}
\end{equation*}
$$

Then, for

$$
\begin{aligned}
& u_{1}^{h}:= \begin{cases}\frac{1}{\left(\varepsilon_{h} / \delta_{h}\right)^{2}}\left(\bar{y}_{1}^{h}-x_{1}\right) & \text { if } r=+\infty, \\
\frac{1}{\varepsilon_{h}}\left(\bar{y}_{1}^{h}-x_{1}\right) & \text { if } r \in\{0,1\},\end{cases} \\
& u_{2}^{h}:= \begin{cases}\frac{1}{\varepsilon_{h} / \delta_{h}}\left(\bar{y}_{2}^{h}-h x_{2}\right) & \text { if } r=+\infty, \\
\frac{1}{\varepsilon_{h} / h}\left(\bar{y}_{2}^{h}-h x_{2}\right) & \text { if } r \in\{0,1\},\end{cases} \\
& u_{3}^{h}:=\frac{1}{\varepsilon_{h} / \delta_{h}}\left(\bar{y}_{3}^{h}-\delta_{h} x_{3}\right), \\
& \vartheta^{h}:=\frac{1}{I_{0}} \frac{1}{h \varepsilon_{h}} \int_{\omega}\left(\delta_{h} x_{2} \bar{y}_{3}^{h}-h x_{3} \bar{y}_{2}^{h}\right) d x_{2} d x_{3},
\end{aligned}
$$

we have that, up to subsequences, $u^{h} \rightarrow \bar{u}$ in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right), \vartheta^{h} \rightharpoonup \bar{\vartheta}$ weakly in $W^{1,2}(0, \ell)$ and $Q^{h} \rightarrow \bar{Q}$, where $(\bar{u}, \bar{\vartheta}, \bar{Q}) \in \mathcal{A}^{r} \times W^{1,2}(0, \ell) \times S O(3)$ minimizes the functional
$J^{r}(u, \vartheta, R):= \begin{cases}I^{r}(u, \vartheta)-R_{32} \int_{0}^{\ell} \bar{f} u_{2} d x_{1}-R_{33} \int_{0}^{\ell} \bar{f} u_{3} d x_{1} & \text { if } r=+\infty, \\ I^{r}(u, \vartheta)-R_{33} \int_{0}^{\ell} \bar{f} u_{3} d x_{1} & \text { if } r \in\{0,1\},\end{cases}$
among all $(u, \vartheta, R) \in \mathcal{A}^{r} \times W^{1,2}(0, \ell) \times S O(3)$. Here we have set $\bar{f}\left(x_{1}\right):=$ $\int_{-\frac{1}{2}}^{\frac{1}{2}} f\left(x_{1}, x_{2}\right) d x_{2}$. Furthermore, we have

$$
\begin{equation*}
\lim _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} \inf J^{h}=\lim _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} J^{h}\left(y^{h}\right)=J^{r}(\bar{u}, \bar{\vartheta}, \bar{Q})=\min J^{r} . \tag{67}
\end{equation*}
$$

Remark 4.2 The two conditions in (61) guarantee that

$$
\int_{\Omega} f^{h} e_{3} \cdot Q x d x=0
$$

for every $Q \in S O(3)$ and every $h>0$.

Proof of Theorem 4.1. Let $\left(y^{h}\right)$ be a minimizing sequence of $\frac{1}{\varepsilon_{h}^{2}} J^{h}$ in the sense of (63) and let $p_{h}(x)=\left(x_{1}, h x_{2}, \delta_{h} x_{3}\right)$ be defined as in (8). Applying Lemma 3.1 with $\eta_{h}^{2}:=I^{h}\left(y^{h}\right)$ and $\eta_{h}$ in place of $\varepsilon_{h}$, and using PoincaréWirtinger inequality we find some constants $P^{h} \in S O(3)$ and $d^{h} \in \mathbb{R}^{3}$ such that for $\tilde{y}^{h}:=P^{h^{T}} y^{h}-d^{h}-p_{h}$ there holds

$$
\begin{equation*}
\left\|\tilde{y}^{h}\right\|_{L^{2}}^{2}+\left\|\nabla_{h} \tilde{y}^{h}\right\|_{L^{2}}^{2} \leq C \delta_{h}^{-2} I^{h}\left(y^{h}\right) \tag{68}
\end{equation*}
$$

Moreover, by (61) we have

$$
\begin{equation*}
\inf J^{h} \leq J^{h}\left(p_{h}\right)=-\int_{\Omega} \delta_{h} x_{3} f^{h}\left(x_{1}, x_{2}\right) d x=0 \tag{69}
\end{equation*}
$$

Using (61) - (63), (68), and (69), we obtain

$$
\begin{aligned}
I^{h}\left(y^{h}\right) & =J^{h}\left(y^{h}\right)+\int_{\Omega} f^{h} P^{h^{T}} e_{3} \cdot \tilde{y}^{h} d x \\
& \leq C \varepsilon_{h}^{2}+\alpha_{h}\left\|\tilde{y}^{h}\right\|_{L^{2}} \leq C \varepsilon_{h}^{2}+C \frac{\alpha_{h}}{\delta_{h}}\left(I^{h}\left(y^{h}\right)\right)^{1 / 2}
\end{aligned}
$$

Recalling that $\varepsilon_{h}=\alpha_{h} / \delta_{h}$, the previous inequality yields

$$
\begin{equation*}
I^{h}\left(y^{h}\right) \leq C \varepsilon_{h}^{2} . \tag{70}
\end{equation*}
$$

Moreover, using again (61), (62), and (68), we deduce

$$
J^{h}\left(y^{h}\right) \geq-\int_{\Omega} f^{h} P^{h^{T}} e_{3} \cdot \tilde{y}^{h} d x \geq-\alpha_{h}\left\|\tilde{y}^{h}\right\|_{L^{2}} \geq-C \varepsilon_{h}^{2}
$$

Since $\left(y^{h}\right)$ is a minimizing sequence, the last inequality together with (69) implies that $\left|\inf J^{h}\right| \leq C \varepsilon_{h}^{2}$.

By (70) and by Theorem 3.11 we deduce statement $i i$ of the theorem.
Assume now (65) and, if $r=+\infty$, (66). By Theorem 3.11 we deduce that there exists $(\bar{u}, \bar{\vartheta}) \in \mathcal{A}^{r} \times W^{1,2}(0, \ell)$ such that, up to subsequences, $u^{h} \rightharpoonup \bar{u}$ weakly in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ (strongly if $r=+\infty$ ) and $\vartheta^{h} \rightharpoonup \bar{\vartheta}$ weakly in $W^{1,2}(0, \ell)$. Moreover, up to subsequences, we also have that $Q^{h}$ converges to some $\bar{Q} \in$ $S O(3)$. By (40) and (62) we obtain that

$$
\begin{align*}
\liminf _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} J^{h}\left(y^{h}\right) & \geq I^{r}(\bar{u}, \bar{\vartheta})+\liminf _{h \rightarrow 0}\left(-\int_{\Omega} \frac{1}{\alpha_{h}} f^{h} Q^{h^{T}} e_{3} \cdot \frac{\delta_{h}}{\varepsilon_{h}}\left(\bar{y}^{h}-p_{h}\right) d x\right) \\
& =J^{r}(\bar{u}, \bar{\vartheta}, \bar{Q}) \tag{71}
\end{align*}
$$

Let now $(u, \vartheta, R) \in \mathcal{A}^{r} \times W^{1,2}(0, \ell) \times S O(3)$. By Theorems 3.12 and 3.14 there exists a sequence $\left(\hat{y}^{h}\right) \subset W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ such that the corresponding displacement $\hat{u}^{h}$ and twist function $\hat{\vartheta}^{h}$ satisfy $\hat{u}^{h} \rightarrow u$ in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right), \hat{\vartheta}^{h} \rightarrow \vartheta$ in $W^{1,2}(0, \ell)$, and

$$
\limsup _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} W\left(\nabla_{h} y^{h}\right) d x \leq I^{r}(u, \vartheta) .
$$

This implies that

$$
\begin{align*}
\limsup _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} J^{h}\left(y^{h}\right) & =\limsup _{h \rightarrow 0}\left(\frac{1}{\varepsilon_{h}^{2}} \inf J^{h}\right) \leq \limsup _{h \rightarrow 0} \frac{1}{\varepsilon_{h}^{2}} J^{h}\left(R \hat{y}^{h}\right) \\
& =\limsup _{h \rightarrow 0}\left(\frac{1}{\varepsilon_{h}^{2}} I^{h}\left(\hat{y}^{h}\right)-\frac{1}{\varepsilon_{h}^{2}} \int_{\Omega} f^{h} R^{T} e_{3} \cdot \hat{y}^{h} d x\right) \\
& \leq I^{r}(u, \vartheta)+\limsup _{h \rightarrow 0}\left(-\int_{\Omega} \frac{1}{\alpha_{h}} f^{h} R^{T} e_{3} \cdot \frac{\delta_{h}}{\varepsilon_{h}}\left(\hat{y}^{h}-p_{h}\right) d x\right) \\
& =J^{r}(u, \vartheta, R) . \tag{72}
\end{align*}
$$

Combining (71) and (72) we deduce the minimality of $(\bar{u}, \bar{\vartheta}, \bar{Q})$ and the convergence of the energies (67).

To conclude it remains to show that $u^{h} \rightarrow \bar{u}$ strongly in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ for $r \in\{0,1\}$. Arguing as in [10, Subsection 7.2], one can infer from (67) that $\frac{1}{\varepsilon_{h}} \operatorname{sym}\left(\nabla_{h} \bar{y}^{h}-I\right)$ converges strongly in $L^{2}\left(\Omega ; \mathbb{R}^{3 \times 3}\right)$. Repeating the proof of part $i i$ of Lemma 3.7, one can show that this implies strong convergence in $W^{1,2}\left(\Omega ; \mathbb{R}^{3}\right)$ of the sequence of displacements $u^{h}$.

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